Measuring Peculiar Motion using Redshift Space Distortion

Large Scale Structure and Galaxy Flows, Quy Nhon July 6th 2016

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Implication of cosmic acceleration

 Breaking down our knowledge of particle physics: we have limited knowledge of particle physics bounded by testable high energy, and our efforts to explain the cosmic acceleration turn out in vain:

Alternative mechanism to generate fine tuned vacuum energy

New unknown energy component

Unification or coupling between dark sectors

 Breaking down our knowledge of gravitational physics: gravitational physics has been tested in solar system scales, and it is yet confirmed at horizon size:

Presence of extra dimension

Non-linear interaction to Einstein equation

• Failure of standard cosmology model: our understanding of the universe is still standing on assumption:

Inhomogeneous models: LTB, back reaction

Implication of cosmic acceleration

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Non-linear interaction to Einstein equation $4\pi G_{\mu\nu} = 4\pi G_{\mu\nu} T_{\mu\nu}$

• Failure of standard cosmology model: our understanding of the universe is still standing on assumption:

Inhomogeneous models: LTB, back reaction

Galaxy clustering seen in redshift space

- Spectroscopy wide surveys have provided the key observables of distance measures and growth functions, such as 2dF, SDSS, WiggleZ, BOSS
- Most unknowns in the universe will be revealed through LSS





Alam et.al 2015; YSS, Koyama 2009

Galaxy clustering seen in redshift space





YSS, Percival 2009



Planned surveys in the future



History and plan for spectroscopy surveys



Galaxy clustering seen in redshift space





YSS, Taruya, Akira 2015

Key cosmological observables of the universe

- Angular diameter distance D_A: Exploiting BAO as standard rulers which measure the angular diameter distance and expansion rate as a function of redshift.
- Radial distance H⁻¹: Exploiting redshift distortions as intrinsic anisotropy to decompose the radial distance represented by the inverse of Hubble rate as a function of redshift.
- Coherent motion G₀: The coherent motion, or flow, of galaxies can be statistically estimated from their effect on the clustering measurements of large redshift surveys, or through the measurement of redshift space distortions.

Standard ruler

D_s ~150 Mpc

$D_s = \Delta z/H(z)$

$D_s = (1+z) D_A(z) \theta$



Anisotropy galaxy clustering



Anisotropy galaxy clustering



Measured anisotropy galaxy clustering



Perpendicular and radial distance measures

 $(D_A, H^{-1}, G_{\delta}, G_{\Theta}, FoG)$



Perpendicular and radial distance measures

 $(D_A, H^{-1}, G_{\delta}, G_{\Theta}, F_{O}G)$



Growth function measurements

$(D_A, H^{-1}, \mathbf{G}_{\delta}, \mathbf{G}_{\Theta}, FoG)$



Growth function measurements





YSS et.al. 2015

Anisotropy correlation without corrections



Improved RSD model



Non-linear corrections

Higher order polynomials

Finger of God effect

$$P_{s}(k,\mu) = P_{gg}(k) + 2\mu^{2}P_{g\theta}(k) + \mu^{4}P_{\theta\theta}(k)$$

$$\begin{split} \mathsf{P}_{\mathsf{s}}(\mathsf{k},\boldsymbol{\mu}) &= [\mathsf{P}_{\mathsf{gg}}(\mathsf{k}) + \underline{\varDelta} \mathsf{P}_{\mathsf{gg}} + 2\boldsymbol{\mu}^2 \mathsf{P}_{\mathsf{g}\Theta}(\mathsf{k}) + \underline{\varDelta} \mathsf{P}_{\mathsf{g}\theta} + \boldsymbol{\mu}^4 \mathsf{P}_{\theta\theta}(\mathsf{k}) + \underline{\varDelta} \mathsf{P}_{\theta\theta}(\mathsf{k}) \\ &+ \boldsymbol{\mu}^2 \mathsf{A}(\mathsf{k}) + \boldsymbol{\mu}^4 \mathsf{B}(\mathsf{k}) + \boldsymbol{\mu}^6 \mathsf{C}(\mathsf{k}) + \dots] \exp[-(\mathsf{k}\boldsymbol{\mu}\boldsymbol{\sigma}_{\mathsf{p}})^2] \end{split}$$

Taruya, Nishimichi, Saito 2010; Taruya, Hiramatsu 2008; Taruya, Bernardeau, Nishimichi 2012

$$\begin{split} \mathsf{P}_{\mathsf{s}}(\mathsf{k}, \boldsymbol{\mu}) &= [\mathsf{Q}_0(\mathsf{k}) + \boldsymbol{\mu}^2 \mathsf{Q}_2(\mathsf{k}) + \boldsymbol{\mu}^4 \mathsf{Q}_4(\mathsf{k}) + \boldsymbol{\mu}^6 \mathsf{Q}_6(\mathsf{k})] \exp[-(\mathsf{k}\boldsymbol{\mu}\boldsymbol{\sigma}_{\mathsf{p}})^2] \\ & \boldsymbol{\xi}(\sigma, \pi) = \int d^3 \mathsf{k} \ \mathsf{P}(\mathsf{k}, \boldsymbol{\mu}) e^{\mathsf{i}\mathsf{k}\mathsf{x}} = \boldsymbol{\Sigma} \boldsymbol{\xi}_{\mathsf{l}}(\mathsf{s}) \ \mathcal{P}_{\mathsf{l}}(\boldsymbol{\nu}) \\ & \boldsymbol{\xi}_{\ell}(\mathsf{s}) = \mathsf{i}^{\mathsf{l}} \int \mathsf{k}^2 \mathsf{d}\mathsf{k} \ \mathsf{P}_{\mathsf{l}}(\mathsf{k}) \ \mathsf{j}_{\mathsf{l}}(\mathsf{k}\mathsf{s}) \end{split}$$

$$\begin{split} \mathsf{P}_0(\mathsf{k}) &= \mathsf{p}_0(\mathsf{k}) \\ \mathsf{P}_2(\mathsf{k}) &= 5/2 \; [3\mathsf{p}_1(\mathsf{k}) - \mathsf{p}_0(\mathsf{k})] \\ \mathsf{P}_4(\mathsf{k}) &= 9/8 \; [35\mathsf{p}_2(\mathsf{k}) - 30\mathsf{p}_1(\mathsf{k}) + 3\mathsf{p}_0(\mathsf{k}) \;] \\ \mathsf{P}_6(\mathsf{k}) &= 13/16 \; [231\mathsf{p}_3(\mathsf{k}) - 315\mathsf{p}_2(\mathsf{k}) - 105\mathsf{p}_1(\mathsf{k}) + 5\mathsf{p}_0(\mathsf{k}) \;] \end{split}$$

 $p_{n}(k) = 1/2 \left[\gamma(n+1/2,\kappa)/\kappa^{n+1/2}Q_{0}(k) + \gamma(n+3/2,\kappa)/\kappa^{n+3/2}Q_{2}(k) + \gamma(n+5/2,\kappa)/\kappa^{n+5/2}Q_{4}(k) + \gamma(n+7/2,\kappa)/\kappa^{n+7/2}Q_{6}(k) + \gamma(n+7/2,\kappa)/$

 $\mathbf{\kappa} = \mathbf{k}^2 \sigma^2_{\mathbf{p}}$

YSS, Okumura, Taruya 2014 Taruya, Nichimishi, Saito 2010



YSS et.al. 2015



YSS et.al. 2015



Open new window to test cosmological models $(D_A, H^{-1}, G_{\delta}, G_{\Theta}, FoG)$ Standard model New physics Cold dark matter Quintessence dark energy Massless neutrino Phantom dark energy

Open new window to test cosmological models

 $(D_A, H^{-1}, G_{\delta}, G_{\Theta}, FoG, New, New, ...)$

Standard model

New physics

Cold dark matter

Massless neutrino

Hot or warm dark matter

Massive neutrino

Interacting dark matter

Unified dark matter

Quintessence dark energy

Phantom dark energy Decaying vacuum Chameleon type gravity Dilaton or Symmetron Vainstein type gravity Inhomogeneity of universe

non-Friedman universe

Open new window to test cosmological models

$(\mathsf{D}_{\mathsf{A}},\,\mathsf{H}^{\text{-1}},\,\mathsf{G}_{\delta},\,\mathsf{G}_{\Theta},\,\mathsf{FoG},\,\,,\,\mathsf{New},\,\ldots)$

New physics

Chameleon type gravity

Probing modified gravity



Probing modified gravity



Probing modified gravity

 $(D_A, H^{-1}, G_{\delta}, G_{\Theta}, FoG, |f_{R0}|)$

We find new constraints on f(R) gravity models using BOSS DR11

If_{R0}I < 8×10⁻⁴ at 95% confidence limit



Open new window to test cosmological models

 $(D_A, H^{-1}, G_{\delta}, G_{\Theta}, FoG, New, \ldots)$

Standard model

Massive neutrino



Probing neutrino mass

We build the neutrino RSD templates at small mass limit $m_v < 1 eV$



Oh, YSS arXiv:1607.01074

Probing neutrino mass



Oh, YSS arXiv:1607.01074

Excavating neutrino mass buried under LSS

The measured neutrino mass is $m_v = 0.2^{+0.28}_{-0.17}$ eV The combined constraints on m_v is distinguishable from massless m_v



Oh, YSS arXiv:1607.01074

Challenge to the future precision cosmology



Mapping of clustering from real to redshift spaces

$$P_{s}(k,\mu) = \int d^{3}x \ e^{ikx} \langle \delta \delta \rangle$$

$$\mathsf{P}_{\mathsf{s}}(\mathsf{k},\mu) = \int d^{3}x \, e^{\mathsf{i}\mathsf{k}x} \langle e^{\mathsf{j}\mathsf{v}} \, (\boldsymbol{\delta} + \mu^{2} \Theta) (\boldsymbol{\delta} + \mu^{2} \Theta) \rangle$$

 $= \int d^3x \ e^{ikx} \exp\{\langle e^{jv} \rangle_c\} \left[\langle e^{jv} (\delta + \mu^2 \Theta) (\delta + \mu^2 \Theta) \rangle_c + \langle e^{jv} (\delta + \mu^2 \Theta) \rangle_c \langle e^{jv} (\delta + \mu^2 \Theta) \rangle_c \right]$

- Higher order polynomials are generated by density and velocity cross-correlation which generate the infinite tower of correlation pairs. We take the perturbative approach to cut off higher orders.
- The FoG effect consists of the one-point contribution and the correlated velocity pair contribution. The latter is perturbatively expanded as F term, and the former is parameterised using σ_p .

Yi, YSS 2016

Direct measurement of higher order polynomials



The residual FoG

Yi, YSS 2016

We subtract out the perturbative higher order polynomials, and the remaining's can be considered to be FoG effect. If our formulation is correct, those all residuals should be consistent in terms of scale, and fitted to be Gaussian with constant σ_p .



Challenge to the precision cosmology



Yi, YSS 2016

Bispectrum in redshift space

 $B(k_1, k_2, k_3, \mu_1, \mu_2) = D^{B}_{FoG} B^{PT}(k_1, k_2, k_3, \mu_1, \mu_2)$



Sabiu, YSS 2016 prepared

YSS, Taruya, Akira 2015

BAO cloud in bispectrum



The full RSD theory for bispectrum

First order (equivalent to Kaiser term)

 $\begin{bmatrix} \langle \Delta \Delta' \Delta'' \rangle_c \\ + j_1 \left(\langle V \Delta \Delta' \Delta'' \rangle_c + \langle \Delta \Delta' \rangle_c \langle V \Delta'' \rangle_c + \langle \Delta'' \Delta \rangle_c \langle V \Delta' \rangle_c + \langle \Delta' \Delta'' \rangle_c \langle V \Delta \rangle_c \right) \\ + j_2 \left(\langle V' \Delta \Delta' \Delta'' \rangle_c + \langle \Delta \Delta' \rangle_c \langle V' \Delta'' \rangle_c + \langle \Delta'' \Delta \rangle_c \langle V' \Delta' \rangle_c + \langle \Delta' \Delta'' \rangle_c \langle V' \Delta \rangle_c \right)$

Second order

$$\begin{split} + j_{1}j_{2}(\langle VV'\Delta\Delta'\Delta''\rangle_{c} + \langle \Delta\Delta'\rangle_{c}\langle VV'\Delta''\rangle_{c} + \langle \Delta''\Delta\rangle_{c}\langle VV'\Delta'\rangle_{c} + \langle \Delta'\Delta''\rangle_{c}\langle VV'\Delta\rangle_{c} \\ + \langle V\Delta\Delta'\rangle_{c}\langle V'\Delta''\rangle_{c} + \langle V'\Delta\Delta'\rangle_{c}\langle V\Delta''\rangle_{c} + \langle V\Delta''\Delta\rangle_{c}\langle V\Delta'\rangle_{c} + \langle V\Delta''\Delta\rangle_{c}\langle V\Delta'\rangle_{c} \\ + \langle V\Delta'\Delta''\rangle_{c}\langle V'\Delta\rangle_{c} + \langle V'\Delta'\Delta''\rangle_{c}\langle V\Delta\rangle_{c} \rangle \\ + j_{1}^{2}(\frac{1}{2}\langle V^{2}\Delta\Delta'\Delta''\rangle_{c} + \frac{1}{2}\langle\Delta\Delta'\rangle_{c}\langle V^{2}\Delta''\rangle_{c} + \frac{1}{2}\langle\Delta''\Delta\rangle_{c}\langle V^{2}\Delta'\rangle_{c} + \frac{1}{2}\langle\Delta'\Delta''\rangle_{c}\langle V^{2}\Delta\rangle_{c} \\ + \langle V\Delta\Delta'\rangle_{c}\langle V\Delta''\rangle_{c} + \langle V\Delta''\Delta\rangle_{c}\langle V\Delta'\rangle_{c} + \langle V\Delta'\Delta''\rangle_{c}\langle V\Delta\rangle_{c} \rangle \\ + j_{2}^{2}(\frac{1}{2}\langle V'^{2}\Delta\Delta'\Delta''\rangle_{c} + \frac{1}{2}\langle\Delta\Delta'\rangle_{c}\langle V'^{2}\Delta''\rangle_{c} + \frac{1}{2}\langle\Delta''\Delta''\rangle_{c}\langle V'^{2}\Delta\rangle_{c} \\ + \langle V\Delta\Delta'\rangle_{c}\langle V\Delta''\rangle_{c} + \langle V\Delta''\Delta\rangle_{c}\langle V'\Delta''\rangle_{c} + \langle V\Delta'\Delta''\rangle_{c}\langle V'\Delta\rangle_{c} \rangle \\ \end{split}$$

FoG term

$$\exp \left\{ \frac{1}{2} (j_1^2 + j_2^2 + j_3^2) \sigma_z^2 - j_1^2 \langle u_z(\vec{r}) u_z(\vec{r}') \rangle_c - j_2^2 \langle u_z(\vec{r}) u_z(\vec{r}'') \rangle_c \right. \\ \left. + j_1 j_2 \left[\langle u_z(\vec{r}') u_z(\vec{r}'') \rangle_c - \langle u_z(\vec{r}) u_z(\vec{r}') \rangle_c - \langle u_z(\vec{r}) u_z(\vec{r}'') \rangle_c \right] \right\}$$

YSS et.al. 2016 prepared

Improved coherent motion in the future



YSS, Taruya, Akira 2015

Conclusion

- We measure coherent motion of the universe with BOSS catalogue using RSD perturbative theory, which provides us with trustable measurements.
- The full perturbative approaches allow us to prove the exotic cosmic acceleration model such as modified gravity of f(R) gravity.
- We probe the non-trivial neutrino mass about 0.2eV, and the measured Hubble constant gets to be even smaller about 65.
- The future experiment opens new precision cosmology era, and we are ready for the challenge. Our new RSD theoretical model is promising to probe coherent motion in a percentage precision.
- The combination of power and bi spectra is essential to probe the coherent motion tightly. We make lots of efforts for it.