

Improved
SLAP code

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Improved lensing reconstruction with a non-parametric code

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Main idea

Based on the SLAP code (Diego et al. 2005) we improve the lensing reconstruction of the mass in a galaxy cluster:

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- ▶ new code is tested with simulated data (strong lensing) that mimics real data

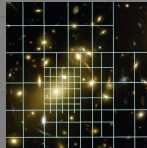
Main idea

Based on the SLAP code (Diego et al. 2005) we improve the lensing reconstruction of the mass in a galaxy cluster:

- ▶ we integrate physical priors in the code: contribution to the deflection field coming from the galaxies of the cluster.
- ▶ new code is tested with simulated data (strong lensing) that mimics real data
- ▶ It greatly help to increase the resolution of the solution and reduce the uncertainties.

SLAP code (Diego et al. 2005)

► Non-parametric code:



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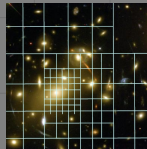
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SLAP code

Results

Conclusions

SLAP code (Diego et al. 2005)

- ▶ Non-parametric code:



- ▶ Approximation: split lens in N_c cells: $M(\theta) = \sum_{i=1}^{N_c} c_i f_i(\theta)$

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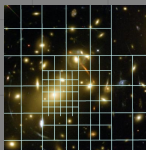
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SLAP code

Results

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$$\alpha_i = \Upsilon_{ij} c_j$$

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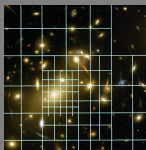
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SLAP code

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- ▶ **The problem:** a linear system of $2N_\theta$ equations and $(2N_S + N_C)$ unknowns:

$$\Theta = \Gamma \mathbf{x} \rightarrow \begin{bmatrix} \theta_x \\ \theta_y \end{bmatrix} = \begin{bmatrix} \Upsilon_x & \mathbf{I}_x & \mathbf{0} \\ \Upsilon_y & \mathbf{0} & \mathbf{I}_y \end{bmatrix} \times \begin{bmatrix} \mathbf{c} \\ \beta_x \\ \beta_y \end{bmatrix}$$

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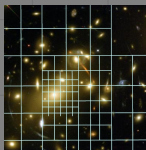
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- ▶ Approximate solution \rightarrow residual equation: $\mathbf{r} \equiv \Theta - \Gamma \mathbf{x}$.

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Improved
SLAP code

Results

Conclusions

SLAP code: analyse & simulate data.

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Results

Conclusions

- ▶ Algorithm for simulations: projects different NFW profiles in the field of view.

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Results

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- ▶ Algorithm for simulations: projects different NFW profiles in the field of view.
- ▶ User defines a number of halos and their characteristics $\{(x, y), M[10^{15} h^{-1} M_{\odot}], r_{soft}, r_s, (e_x, e_y, e_z)\}$

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- We simulate a cluster similar to *A1689* to then check how works the improved version of the code.

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Improved
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- We will use these tool to obtain the deflection field of the cluster galaxies.

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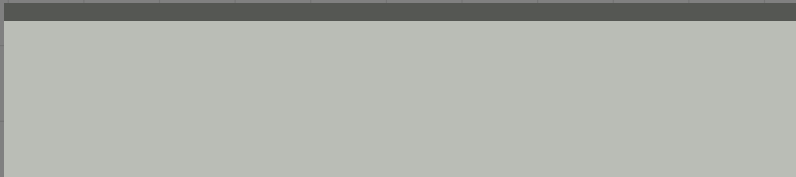
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Improved
SLAP code

Results

Conclusions



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- ▶ Cluster galaxies contribute with a mass proportional to a fiducial value whose proportionality constant is later inferred as a part of the method.

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Improved
SLAP code

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SLAP code

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Improved
SLAP code

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Improved
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- ▶ New unknown, c_{gal} , proportionality constant.
- ▶ **New system of equations:**

$$\begin{bmatrix} \theta_x \\ \theta_y \end{bmatrix} = \begin{bmatrix} \Upsilon_x & \alpha_{gal,x} & \mathbf{I}_x & \mathbf{0} \\ \Upsilon_y & \alpha_{gal,y} & \mathbf{0} & \mathbf{I}_y \end{bmatrix} \times \begin{bmatrix} \mathbf{c} \\ c_{gal} \\ \beta_x \\ \beta_y \end{bmatrix}$$

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Improved
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Conclusions

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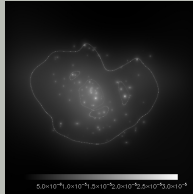
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Improved
SLAP code

Results

Conclusions

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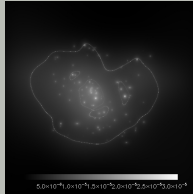
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Improved
SLAP code

Results

Conclusions

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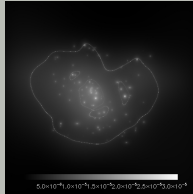
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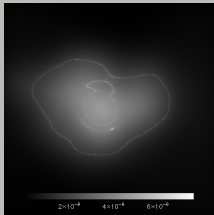
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Results

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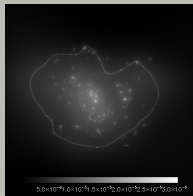
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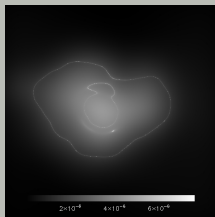
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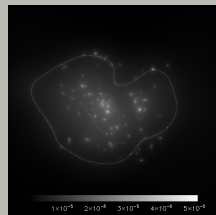
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Results

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Results

Conclusions

Recomposed mass profile

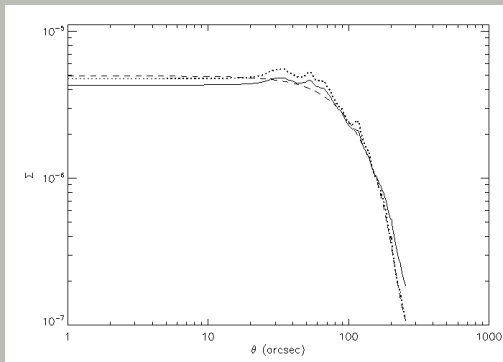


Figure: **Thick lines** → real/simulated case, **dotted lines** → new version of SLAP. and **dashed lines** → old version of SLAP

Conclusions

- ▶ **The addition of cluster galaxies deflection map greatly improves the overall results, giving a reconstructed mass distribution closer to the real one with the same sub-structures.**

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Improved
SLAP code

Results

Conclusions

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Improved
SLAP code

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- ▶ **The addition of cluster galaxies deflection map greatly improves the overall results, giving a reconstructed mass distribution closer to the real one with the same sub-structures.**
- ▶ **Next step:** Apply this new version of the code to real clusters.
- ▶ **This new version of the code based in a non-parametric approach will also provide an important consistency check for the parametric approach, since concurring results will strengthen their validity, whereas any resulting differences would need to be addressed.**

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Improved
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Results

Conclusions