Workshop Fundamental Issues of the Standard Cosmological Model, Cargese



Nontopological solitons: explicit solutions and stability issues E.Nugaev, (INR RAS)

Localized field configuration (soliton) \rightarrow DM candidate

motivation:

- to avoid restrictions on WIMP,
- possible objects in the new field content,
- baryogenesis, cogenesis,
- new mechanism of production?

main issues:

- stability
- energy spectrum
- production in Early Universe

Stability Charge or Topology

 $\begin{array}{l} \mbox{Static solutions} \rightarrow \mbox{problem with Derrick theorem} \\ \mbox{(nonlinear kinetic term, gauge fields)} \\ \mbox{For pure scalar field theory scaling arguments restrict number of space-time} \\ \mbox{dimensions } D < 3 \end{array}$

Stationary solution for U(1)-invariant scalar field theory: $\phi = e^{i\omega t}f(r)$ in ordinary (3+1) space-time (we also turn off gravity) $\mathcal{L} = \partial_{\mu}\phi^*\partial^{\mu}\phi - V(\phi^*\phi)$ only r dependence \rightarrow Q-ball Charge (not electric!) \rightarrow global U(1) symmetry,

G. Rosen, J. Math. Phys. 9 (1968) 996

or Q-balls, S.R. Coleman, Nucl. Phys. B 262 (1985) 263 [Erratum 269 (1986) 744]

overshoot-undershoot method, where r corresponds to time in Eq. of motion

Eq. is nonlinear, how to check (numerical) result?

$$\partial E/\partial Q = \omega$$

here

$$E = \int d^3x (\partial_0 \phi^* \partial_0 \phi + \partial_i \phi^* \partial_i \phi + V),$$

and $Q = i \int d^3x (\partial_0 \phi^* \phi - \phi^* \partial_0 \phi)$

with simple generalisation for additional mediator scalar field \rightarrow NONTOPOLOGICAL SOLITONS

R.Friedberg, T.D. Lee, A.Sirlin, PRD 13(1976) 2739

Time-dependent background: how to investigate stability?

$$V = m^2 \phi^* \phi - \lambda (\phi^* \phi)^2$$

D.L.T. Anderson, G.H. Derrick, J. Math. Phys. 11 (1970) 1336

all solutions are unstable,

$$\delta \phi = e^{i\omega t + |\gamma|t} (\psi_1(r) + \psi_2^*(r))$$

$$\begin{pmatrix} \hat{O}_{11} & \hat{O}_{12} \\ \hat{O}_{21} & \hat{O}_{22} \end{pmatrix} \begin{pmatrix} \psi_1 \\ \psi_2 \end{pmatrix} = \begin{pmatrix} (\omega - \gamma)^2 \psi_1 \\ (\omega + \gamma)^2 \psi_2 \end{pmatrix}$$

Are the stable solutions in any other theory?

- Yes

Q-criterion of stability: $\frac{dQ}{d\omega} < 0$

N.G. Vakhitov, A.A. Kolokolov, Radiophys. Quantum Electron. 16 (1973) 783 -NSE

R.Friedberg, T.D. Lee, A.Sirlin, PRD 13(1976) 2739



R.Friedberg, T.D. Lee, A.Sirlin, PRD **13**(1976) 2739; M.G. Alford, Nucl. Phys. B **298** (1988) 323

Explicit solution: Parabolic-piecewise potential (it was in original Rosen's paper!)

$$V(\phi^*\phi) = M^2 \phi^* \phi \,\theta \left(1 - \frac{\phi^*\phi}{v^2}\right) + (m^2 \phi^* \phi + \Lambda) \theta \left(\frac{\phi^*\phi}{v^2} - 1\right)$$



In this approach $\hat{O}_{12}, \hat{O}_{21}$ are proportional to $\delta(\phi^*\phi - v^2) \rightarrow \delta(r - R)$, where R: f(R) = v and problem for linear excitations can be solved, although full operator is not hermitian! Q-criterion was confirmed.

I. Gulamov, E.N., M. Smolyakov, PRD 87 (2013) 085043

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For flat potential E(Q) is similar to previous works. Spectrum of classically stable solutions starts from cusp point and crosses condensate line at some Q_S . As a result, these solution has no main particle-like feature:

How to make $\delta Q \ll Q$

Indeed, for monopole we have precisely one value of the mass! Let us consider gauge case! Coulomb potential enlarge energy for large Q.

K.-M. Lee, J.A. Stein-Schabes, R. Watkins and L.M. Widrow, PRD 39 (1989) 1665

Perturbation theory works well till $Q \ll e^4$. Check $\partial E/\partial Q = \omega < M$ also holds and this is a obstacle for the second cross of condensate line from below.

I. Gulamov, E.N., M. Smolyakov, PRD 89 (2014) 085006

In weak coupling regime $e \ll 1$ difference between Q_S and Q_{Max} still large.



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Thus, particle-like spectrum for Q-ball can be obtained in potential with turnover (or false vacuum at $\phi = 0$). E.N., M. Smolyakov, JHEP 1407 (2014) 009

At large values one can cure potential by growing terms. This procedure does not disturb Q-ball, which has relatively small field value in the core.

What we can learn from the analogy with monopole?

Are string-like objects are possible?

- Yes, M. Volkov, E Wohnert. Spinning Q balls. PRD66:085003,2002.

 $\phi = e^{i\omega t} e^{in\theta} f(\rho)$

here ρ is two-dimensional radius.

more interesting analogy in classical mechanics

What about stability:

- in (2+1)-dimensional theory
- in our space-time

We have tools to investigate classical stability in (2 + 1) dimensions – just more matching points and Bessel functions. The same check $\frac{dE}{d\omega} = \omega \frac{dQ}{d\omega}$ holds for values per unit length.

Deviation from Q-criterion was found n > 0. E.N., A. Shkerin, PRD 90 (2014) 016002 - Only n = 0 solution is stable in (2 + 1)-dimensions.

This is similar to the Abrikosov vortices in superconductor. (3 + 1)-dimensional case. No additional windings \rightarrow Coleman analysis. The chain of Q-balls is energetically more preferable.

Conclusions

- Parabolic-pieswise potential allows us to find analitic solution
- Meanwhile, general check $\frac{dE}{d\omega} = \omega \frac{dQ}{d\omega}$ is useful
- Moreover, stability issue can be reformulated as algebraic problem
- There is a possibility to tune E and Q of solutions in parametrically small region
- In such a case the spectra of stable Q-balls look similar to the one of free particles

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THANK YOU!