## Bubble Observers in Bubbland

Local Observables in a Landscape of Infrared Gauge Modes


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## Bubble Zero

i Inflation, AKA Bubbland
ii Vectors \& Anisotropies
iii IR Fluctuations \& Bias
iv Background Precession
v Results

M Thursrud, D Mota, F Urban, arXiv:1311.3302
M Thursrud, F Urban, D Mota, arXiv:1312.7491

Bubble One


## Bubble Two

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- Now: quantum fluctuations are a statistical object...

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$$
\Longrightarrow \quad \overrightarrow{\mathcal{E}}_{\mathbb{R}}(\eta)=\int_{\text {damono of ine }}^{\mathcal{H}} d^{3} k e^{-i \vec{k} \vec{x}} \delta \overrightarrow{\mathcal{E}}(\vec{k})
$$

## Infrared Statistics - Single Vector

- We are limited UV observers, so we do not directly probe $\overrightarrow{\mathcal{E}}_{\mathrm{IR}}$
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## Spectrum

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One parameter: Amplitude $g(k) \sim-\left|\mathcal{E}_{\mathrm{IR}}\left(\eta_{0}\right)\right|^{2} \mathcal{N}_{k}^{2}$

## Probability Distributions



## Precession

$$
\text { Tweedledum }\left(k_{1}\right) \text { and Tweedledee }\left(k_{2}\right)
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## Tweedledum ( $k_{1}$ ) and Tweedledee ( $k_{2}$ )

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$\Rightarrow \quad \overrightarrow{\mathcal{E}}_{\mathrm{IR}}$ will be pointing in a different direction


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The background vector makes a random walk in direction space




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Poor little thing...

## Infrared Statistics - Multiple Vector

Multiple vector case - identical coupling: $\mathcal{L}=\sum_{a} f^{2}(\varphi) F_{\mu \nu}^{a} F_{a}^{\mu \nu}$

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Spectrum

$$
\mathcal{P}_{\zeta}(\vec{k})=\mathcal{P}_{\zeta}^{0}(k)\left[1+g(k)\left(\mathcal{A}_{\hat{k}} \cos \chi+\mathcal{B}_{\hat{k}} \sin \chi\right)\right]
$$

$$
\mathcal{A}_{\hat{k}} \sim 3 \cos ^{2} \vartheta-1, \quad \mathcal{B}_{\hat{k}} \sim \sin 2 \vartheta \cos \varphi
$$

Two parameters: Amplitude $g(k)$, Shape $\chi$

## Probability Distributions I



## Probability Distributions II



## How Likely Are We?



## Summary

- Inflation generates Bubbland
- Bubbland is comprised of a multitude of Bubbles
- As observers, we have access to only one Bubble
- Our link with "The Theory" is statistical - observations are biassed
- Vectors generate anisotropies
- Spectator gauge fields can develop into a classical vector background
- Curvature perturbations are quadrupole-modulated: $\mathcal{P}_{\zeta}^{0}(k)\left[1+g(k) \cos ^{2} \vartheta\right]$
- (Non-)Observations of $g(k)$ put statistical constraints on $\mathcal{N}$
- The precession effect
- The background vector is not a constant, but precesses with time
- In the multi-vector case this produces two important features:
a. The quadrupole amplitude $g(k)$ can be positive
b. We need one further shape parameter $\chi$ to describe the correction:

$$
\mathcal{P}_{\zeta}^{0}(k)\left[1+g(k)\left(\mathcal{A}_{\hat{k}} \cos \chi+\mathcal{B}_{\hat{k}} \sin \chi\right)\right]
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$\Rightarrow$ All in all, living in the Bubble can be quite deceiving $\leftarrow$ bottom line

