

# UV renormalization, locality and global hyperbolicity

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Usually in QFT one assumes that

- UV phenomena are local
- UV renormalization can be done via analytical continuation from Euclidian to Minkowskian signature
- QFT in curved space-time is full of surprises:  
dS — in IR, AdS — in UV.

# UV renormalization in $x$ -space

- We consider a scalar field theory:

$$S = \int d^4x \sqrt{-g} \left[ \frac{1}{2} (\partial_\mu \phi)^2 - \frac{1}{2} m^2 \phi^2 - \frac{\lambda}{4!} \phi^4 \right].$$

- The one loop contribution to the effective action of the theory:

$$\Gamma^{(4)} = -i \frac{3\lambda^2}{2(4\pi^2)^2} \int d^4x \int d^4y \phi^2(x) \phi^2(y) F^2(x-y).$$

- Here

$$F(x) \approx \frac{1}{4\pi^2} \frac{i}{x^2 - i\epsilon}$$

is **the most singular part** of the Feynman propagator in position space when  $x^2 \rightarrow 0$ .

# UV renormalization in $x$ -space

- We may extract the leading divergent contribution by changing the variables  $x^\mu = X^\mu + \frac{z^\mu}{2}$ ,  $y^\mu = X^\mu - \frac{z^\mu}{2}$ ,  $\mu = 0, 1, 2, 3$  and by diagonally expanding  $\phi^2(X + z/2) \phi^2(X - z/2)$
- In fact,

$$\Gamma^{(4)} = -i \frac{3\lambda^2}{2(4\pi^2)^2} \int d^4X \phi^4(X) \int d^4z \frac{1}{(z^2 - i\epsilon)^2} + \text{finite terms}$$

The  $z$ -integral provides the standard logarithmic UV divergence.

- Note the importance of the proper  $i\epsilon$  prescription. If one replaces the **Feynman** propagator with the **Wightman** function, then:

$$\int d^4z \frac{1}{\left[(z_0 - i\epsilon)^2 - \vec{z}^2\right]^2} = 0.$$

# A simple example

- We continue by considering the above theory in flat space-time but in the presence of an ideal mirror placed at  $x_3 = 0$ . The ideal mirror reflects all the modes equally well, irrespectively of their momenta. This is expressed by the boundary condition  $\phi|_{x_3=0} = 0$  at the mirror.
- A real physical mirror is definitely transparent to very high energy modes. On general physical grounds one can expect that, if  $a$  is a characteristic interatomic distance of the material of the mirror, a mode whose wavelength  $k$  is much larger than  $1/a$  will not see the mirror at all.
- A real physical mirror can be modeled by a potential barrier which reflects some of the modes and is transparent to the other ones, e.g.:

$$[\square + m^2] \phi = \alpha \delta(x_3) \phi.$$

# A simple example

- The most singular part of the Feynman propagator in presence of an ideal mirror is the following distribution

$$F_{mir}(x, y) \approx \frac{i}{4\pi^2} \frac{1}{s - i\epsilon} - \frac{i}{4\pi^2} \frac{1}{\bar{s} - i\epsilon},$$

where  $s = (x - y)^2$  and  $\bar{s} = (x - \bar{y})^2$  and  $\bar{y}$  is the mirror image of the source point  $y$ .

- In Euclidean signature,  $s$  vanishes only when  $x = y$  and  $\bar{s}$  only when  $x = \bar{y}$ . But the point  $\bar{y}$  does not belong to the portion of space-time that we are considering,  $x_3 > 0$  and  $y_3 > 0$ . Hence, inside the loops in Euclidean signature  $\bar{y}$  plays no role.
- In Lorentzian signature,  $s$  and  $\bar{s}$  vanish on the light-cones whose tips are  $y$  and, respectively,  $\bar{y}$ . Therefore, even though  $\bar{y}$  does not belong to the space-time manifold its light-cone penetrates into it.

# A simple example

- The first singularity in  $F_{mir}(x, y)$  provides the same contribution as in empty space. As regards the second term we have to do the following diagonal expansion:

$$x^\mu - \bar{y}^\mu = z^\mu, x^\mu + \bar{y}^\mu = 2X^\mu, \mu = 0, 1, 2, 3.$$

- Then, the effective action contains the following term:

$$\Gamma_{\bar{s}=0}^{(4)} = -i \frac{3\lambda^2}{2(4\pi^2)^2} \int d^4X \int_{z_3 \geq |X_3|} d^4z \frac{\phi^2(X + z/2) \phi^2(\bar{X} - \bar{z}/2)}{(z^2 - i\epsilon)^2}.$$

- Even though  $z^2$  may vanish the components of four-vector  $z$  are generically not small and the diagonal expansion of  $\phi^2(X + z/2) \phi^2(\bar{X} - \bar{z}/2)$  cannot be performed.
- Both the singularities of  $F_{mir}(x, y)$  at  $s = 0$  and  $\bar{s} = 0$  do contribute to the UV divergence of the integral on the right hand side, while the mixed terms contribute finite expressions.

# The geometry of AdS and Lobachevsky spaces

- The 4-dimensional **Euclidean AdS (Lobachevsky space)** space

$$EAdS_4 := \{X_0^2 - X_1^2 - X_2^2 - X_3^2 - X_4^2 = 1\},$$

is one of the sheets (say  $X_0 \geq 1$ ) of the two-sheeted real hyperboloid embedded into:

$$ds^2 = dX_0^2 - dX_1^2 - dX_2^2 - dX_3^2 - dX_4^2.$$

- The 4-dimensional **(Lorentzian) AdS space**:

$$AdS_4 := \{X_0^2 - X_1^2 - X_2^2 - X_3^2 + X_4^2 = 1\}.$$

embedded into

$$ds^2 = dX_0^2 - dX_1^2 - dX_2^2 - dX_3^2 + dX_4^2.$$

- The EAdS and AdS are related to each other via the analytic continuation  $X_4 \rightarrow i X_4$ . We set the curvatures of the hypeboloids to one:  $R = 1$ .

# The geometry of AdS and Lobachevsky spaces

- The hyperbolic distance is defined via the **invariant scalar product**:

$$\xi = \eta_{\mu\nu} X^\mu Y^\nu = \cosh d(X, Y),$$

where  $d(X, Y)$  is the geodesic distance.

- In **EAdS** the hyperbolic distance is  $\xi \geq 1$ , because  $d(X, Y)$  is real, while in **AdS**  $\xi$  can take any real value.
- The Feynman propagator obeys

$$\begin{aligned} [\square + m^2] F(X, Y) &= [(1 - \xi^2) \partial_\xi^2 - 4\xi \partial_\xi + m^2] F(\xi) = \\ &= 4\pi \delta(X, Y) + 4\pi i e^{-i\pi\nu} \delta(X, -Y). \end{aligned}$$

# The Feynman propagator in AdS and Lobachevsky spaces

- The Feynman propagator can be represented both in **EAdS**, in **global AdS** and in its covering  $\widetilde{\text{AdS}}$  manifold as follows:

$$F(X, Y) = A_+ {}_1F_2\left(\frac{3}{2} - \nu, \frac{3}{2} + \nu; 2; \frac{1 + \xi + i\epsilon}{2}\right) + \\ + A_- {}_1F_2\left(\frac{3}{2} - \nu, \frac{3}{2} + \nu; 2; \frac{1 - \xi - i\epsilon}{2}\right)$$

- $\nu = \sqrt{\frac{9}{4} + m^2}$
- When  $\xi^2 \rightarrow 1$  there is the following leading singularities of the AdS Feynman propagator:

$$F(\xi) \approx -\frac{i}{8\pi^2 (\xi - 1 + i\epsilon)} - \frac{e^{-i\pi\nu}}{8\pi^2 (\xi + 1 + i\epsilon)}.$$

# The Feynman propagator in AdS and Lobachevsky spaces

- The first singularity, at  $\xi = 1$ , is the same as in flat empty space. Note that:

$$(X - Y)^2 - i\epsilon = 2(1 - (\xi + i\epsilon))$$

- The second singularity, at  $\xi = -1$ , is when  $X$  sits on the light cone with the apex at  $\tilde{Y} = -Y$  — point antipodal to  $Y$ .
- In **Lobachevsky space** the second singularity is not seen, because there  $\xi \geq 1$ . But in **AdS** the second singularity is present.

# Loop corrections in AdS and Lobachevsky spaces

- In global AdS the relevant part of the correction is:

$$\Gamma^{(4)} \propto \lambda^2 \int d^5 X \delta(X^2 - 1) \int d^5 Y \delta(Y^2 - 1) \phi^2(X) \phi^2(Y) \times \\ \times \left[ \frac{1}{(X - Y)^2 - i\epsilon} - \frac{1}{(X + Y)^2 - i\epsilon} \right]^2.$$

- The first pole leads to the same renormalization as in flat spacetime. The second pole is different and leads to divergences of a new type. The cross terms lead to less singular contributions.
- Thus, we have to introduce a new counter-term into the Lagrangian:

$$\Delta\mathcal{L} = \frac{\gamma e^{-2\pi i\nu}}{4} \phi^2(X) \phi^2(-X), \quad (1)$$

with a complex coefficient depending on the mass parameter and a new coupling constant  $\gamma$ .