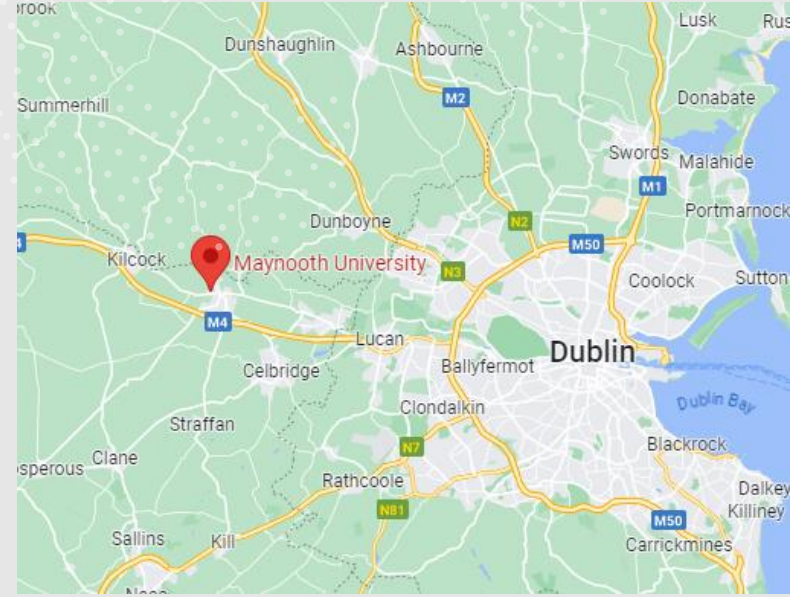


Evolution of Cosmic Voids in the Schrödinger- Poisson Formalism

Aoibhinn Gallagher
Maynooth University



**Maynooth
University**
National University
of Ireland Maynooth

Schrödinger-Poisson System

- Applications in fuzzy dark matter
- Much quicker and simpler to compute

$$\rho = \psi\psi^* = \alpha^2$$

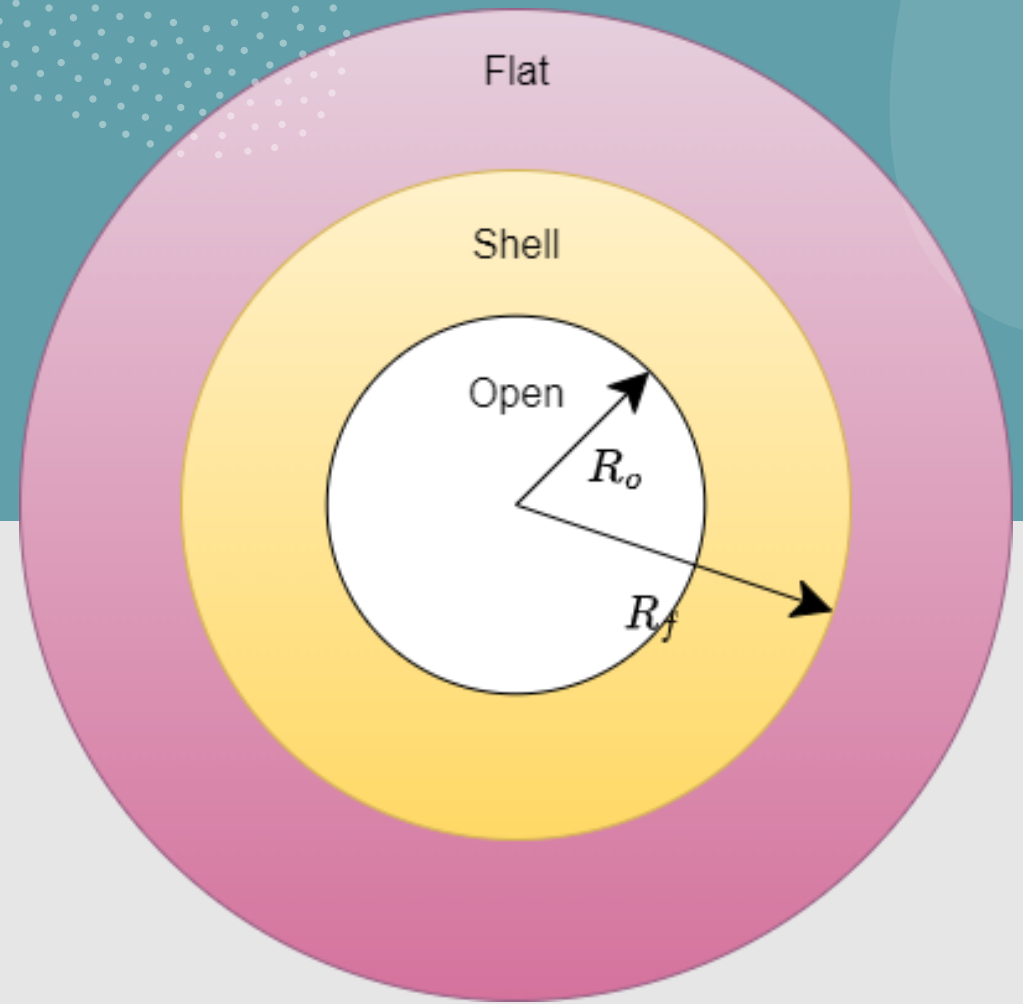
$$\psi = \alpha e^{i\phi/\nu}$$

$$P = \frac{\nu^2}{2} \frac{\nabla^2 \alpha}{\alpha}$$

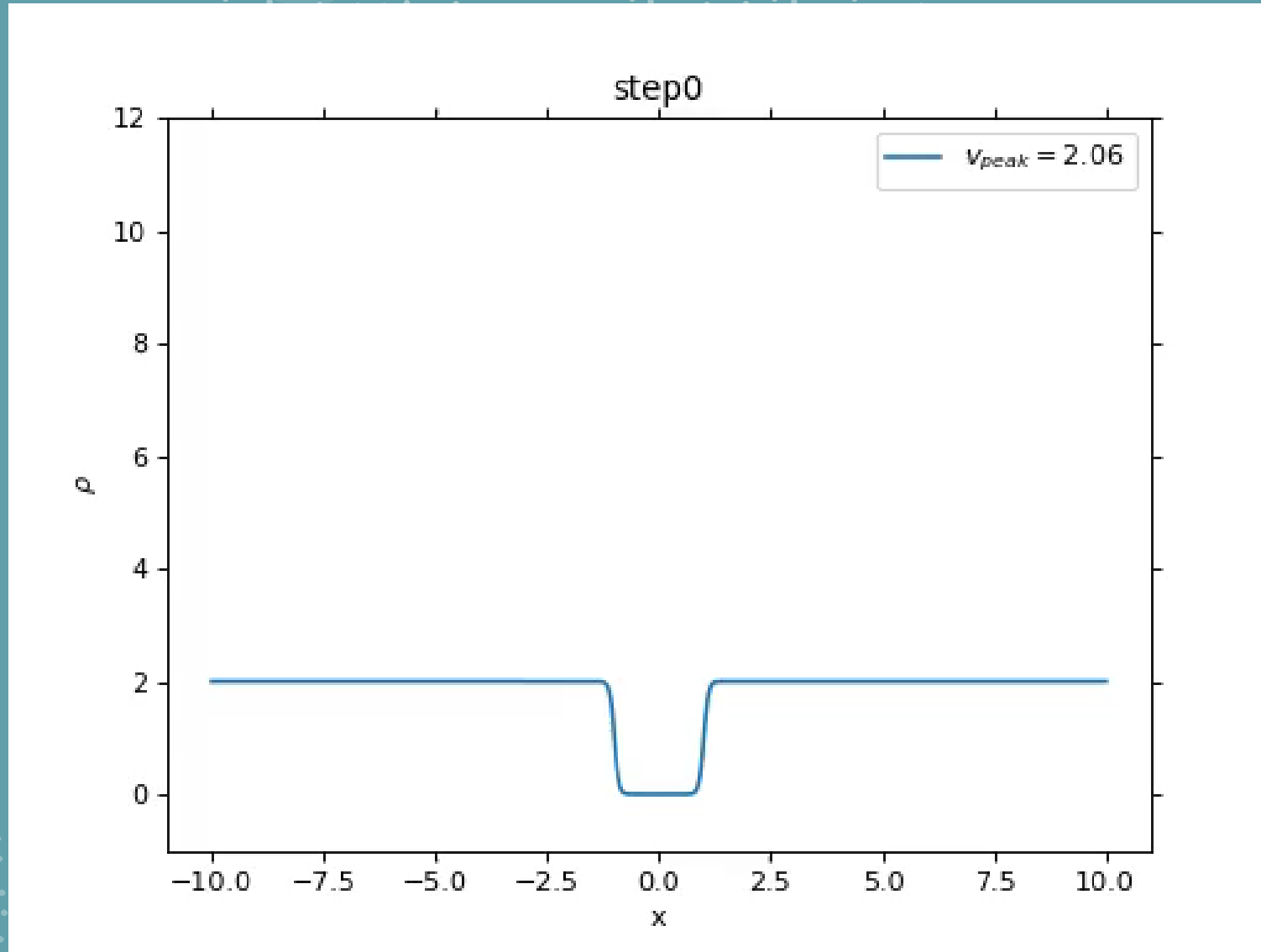
$$i\nu \frac{\partial \psi}{\partial t} = -\frac{\nu^2}{2} \nabla^2 \psi + V\psi + P\psi$$

$$\nabla^2 V = 4\pi G |\psi|^2$$

Modelling Voids



Void Evolution



What's next?

- Getting a SP type system starting from Navier-Stokes
 - Entropy of a Schrödinger fluid
- Using Schrödinger- Poisson for Reconstruction

Back-up slide

