

Quantum Cosmological Perturbations of Multiple Fluids

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Workshop Fundamental Issues of the Standard Cosmological Model

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Outline

Motivation

- Find a consistent way to perform the quantization of coupled multiple components.
 - Multifield inflation.
 - Multiple fluids in the contracting branch of a bouncing model.
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- Obtain the Hamiltonian for multiple-fluid perturbations plus metric perturbations.
 - The only interaction considered between fluids is through gravity.
 - Solving the Hamiltonian and momentum constraints leads naturally to gauge invariant variables.
 - General quantization in a non-canonical form.
 - Different field representations leads to (non)unitary evolution.
 - Trivial field representation.
 - Two fluids example.

Metric perturbation

Given the background metric $\bar{g}_{\mu\nu}$ and the perturbed metric $g_{\mu\nu}$ we define the perturbation tensor

$$\xi_{\mu\nu} \equiv g_{\mu\nu} - \bar{g}_{\mu\nu},$$

and the difference connection tensor

$$\begin{aligned} (\nabla_\mu - \bar{\nabla}_\mu) A_\nu &= \mathcal{F}_{\mu\nu}{}^\beta A_\beta, \\ \mathcal{F}_{\alpha\beta\gamma} &= -\frac{1}{2} (\bar{\nabla}_\alpha \xi_{\beta\gamma} + \bar{\nabla}_\beta \xi_{\gamma\alpha} - \bar{\nabla}_\gamma \xi_{\alpha\beta}). \end{aligned}$$

- The use of the above variables greatly simplifies the calculations.
- The above variables are true tensors.
- Bar variables represent background quantities, all definitions are valid for both background and perturbed variables.

Second Order Lagrangian

Using the introduced variables we obtain the following second order Lagrangian,

$$\begin{aligned}\delta\mathcal{L}_{gk}^{(2)} &= \frac{\sqrt{-\bar{g}}}{2\kappa} \left[\mathcal{F}^{\mu\nu\gamma} \mathcal{F}_{\gamma(\mu\nu)} - \mathcal{F}_{a\mu} \mathcal{F}_b{}^\mu \right], \\ \delta\mathcal{L}_{gP}^{(2)} &= \frac{\sqrt{-\bar{g}}}{2\kappa} \left(\bar{G}_{\mu\nu} + \frac{\bar{g}_{\mu\nu}}{4} \bar{R} \right) \xi^\mu{}_\alpha \left(\xi^{\alpha\nu} - \frac{\bar{g}^{\alpha\nu} \xi}{2} \right),\end{aligned}$$

where we have discarded the surface terms and $\kappa = 8\pi G/c^4$.

Background Foliation

To introduce a time direction we define a background foliation by the vector geodetic field $\bar{v}^\mu$. The same foliation is used to split the perturbed metric, however, the field must be normalized with respect to $g_{\mu\nu}$ giving v^μ . We define the time derivative by

$$\dot{T} = \bar{\gamma} [\mathcal{L}_{\bar{v}} T].$$

The hypersurfaces given by v^μ introduce:

- the projector $\gamma_{\mu\nu} = g_{\mu\nu} + v_\mu v_\nu$, and operator $\bar{\gamma}[A_\mu] = \bar{\gamma}_\mu{}^\nu A_\nu$,
- the extrinsic curvature $\nabla_\mu v_\nu = \mathcal{K}_{\mu\nu}$,
- the Expansion factor $\Theta = \mathcal{K}_\mu{}^\mu$ and shear $\sigma_{\mu\nu} \equiv \mathcal{K}_{\mu\nu} - \frac{\Theta}{3}\gamma_{\mu\nu}$,
- the unique covariant derivative D_μ compatible with $\gamma_{\mu\nu}$,
- the spatial curvature $\mathcal{R}_{\mu\nu\alpha}{}^\beta A_\beta \equiv [D_\mu D_\nu - D_\nu D_\mu]A_\alpha$.

Background Projection

With the background foliation we write the tensor $\xi_{\mu\nu}$ as

$$\xi_{\mu\nu} = 2\phi\bar{v}_\mu\bar{v}_\nu + 2B_{(\mu}\bar{v}_{\nu)} + 2C_{\mu\nu}, \quad B_\mu\bar{v}^\mu = C_{\mu\nu}\bar{v}^\mu = 0.$$

And the scalar vector tensor (SVT) decomposition

$$\begin{aligned} B_\mu &= \bar{D}_\mu \mathcal{B} + \mathcal{B}_\mu, \\ C_{\mu\nu} &= \psi\gamma_{\mu\nu} - \bar{D}_\mu \bar{D}_\nu \mathcal{E} + \bar{D}_{(\mu} \mathcal{F}_{\nu)} + W_{\mu\nu}, \\ C_{\mu\nu} &= \psi^{\dagger}\gamma_{\mu\nu} - \bar{D}_{\langle\mu} \bar{D}_{\nu\rangle} \mathcal{E} + \bar{D}_{(\mu} \mathcal{F}_{\nu)} + W_{\mu\nu}, \end{aligned}$$

where $\bar{D}_\mu \mathcal{B}^\mu = \bar{D}_\mu \mathcal{F}^\mu = \bar{D}_\nu W_\mu{}^\nu = W_\mu{}^\mu = 0$.

Kinematic Perturbations

The advantage of the kinematic variables is that they are directly related to the tensors $\mathcal{F}_{\mu\nu}{}^\gamma$ and $\xi_{\mu\nu}$. For example,

$$\bar{\gamma} [\delta \mathcal{K}_{\mu\nu}] = -\phi \bar{\mathcal{K}}_{\mu\nu} + \bar{\gamma} [\mathcal{F}_{\mu\nu}{}^\gamma \bar{v}_\gamma].$$

Substituting in the Lagrangian one obtains

$$\begin{aligned} \delta \mathcal{L}_g^{(2)} = \frac{\sqrt{-\bar{g}}}{2\kappa} & \left\{ \delta \mathcal{K}_\mu{}^\nu \delta \mathcal{K}_\nu{}^\mu - \delta \Theta^2 - C_\mu{}^\nu \delta \mathcal{R}_\nu{}^\mu + \left(\frac{C}{2} - \phi \right) \delta \mathcal{R} \right. \\ & + 4C^{\mu\sigma} \bar{\sigma}_\sigma{}^\gamma B_{[\mu||\gamma]} + 2 (C_{\mu\nu||\gamma} - C_{||\mu} \bar{\gamma}_{\nu\gamma}) \bar{\sigma}^{\mu\nu} B^\gamma \\ & + 4C_\sigma{}^\lambda C_\beta{}^\alpha (\bar{\sigma}^{\sigma\beta} \bar{\sigma}_{\lambda\alpha} - \bar{\gamma}^{\sigma\beta} \bar{\sigma}_\lambda{}^\gamma \bar{\sigma}_{\gamma\alpha}) \\ & + \left[(\bar{\Theta}\phi + B^\gamma{}_{||\gamma} - \delta\Theta) C_\mu{}^\nu - \left(\phi \bar{\mathcal{K}}_\nu{}^\mu + B_\mu{}^{||\nu} - \delta \mathcal{K}_\mu{}^\nu \right) C \right] \bar{\sigma}_\nu{}^\mu \\ & + \bar{G}_{\bar{v}\bar{v}} (B_\mu B^\mu - \phi^2 - 2\phi C) + \bar{G}_{\bar{v}\mu} (4B_\alpha C^{\alpha\mu} - 2CB^\mu) \\ & \left. + \bar{G}_{\mu\nu} (2C^{\mu\alpha} C_\alpha{}^\nu - C^{\mu\nu} C) \right\}. \end{aligned}$$

Action Principle for a Thermodynamic Fluid

In the action principle for a thermodynamic fluid,¹ we write the specific enthalpy as the norm of a field, i.e., $\vartheta \equiv \sqrt{-\vartheta_\mu \vartheta_\nu g^{\mu\nu}}$. The field is defined by the following potentials and specific entropy

$\vartheta_\mu = \nabla_\mu \varphi_1 + \varphi_2 \nabla_\mu \varphi_3 + \varphi_4 \nabla_\mu s$. The action for the fluid is

$$S_m = \int d^4x \mathcal{L}_m, \quad \text{with} \quad \mathcal{L}_m = \sqrt{-g} p(\vartheta, s).$$

It is useful to express the matter field perturbations in terms of the velocity potential,

$$u_\mu = \bar{\gamma} \left[\frac{\delta \vartheta_\mu}{\bar{\vartheta}} \right] = \mathcal{V}_{\parallel\mu}, \quad \mathcal{V} \equiv \frac{\delta \varphi_1 + \bar{\varphi}_2 \delta \varphi_3 + \bar{\varphi}_4 \delta s}{\bar{\vartheta}},$$

and the gauge invariant variable

$$\Pi_{\mathcal{V}} = -\sqrt{-\bar{g}} \left(\delta \rho - \bar{\Theta}(\bar{\rho} + \bar{p}) \mathcal{V} \right).$$

¹B. Schutz, *Phys. Rev. D* 2.12 (1970), 2762.

Thermodynamic Fluid Lagrangian

Each fluid contributes with a Lagrangian in the form:

$$\frac{\delta \mathcal{L}_{mi}^{(2)}}{\sqrt{-g}} = \frac{\bar{c}_{si}^2 \Pi_{\mathcal{V}_i}^2}{2\sqrt{-\bar{g}}^2 (\bar{\rho}_i + \bar{p}_i)} + \frac{(\bar{\rho}_i + \bar{p}_i)}{2} \mathcal{V}_i \bar{D}_K^2 \mathcal{V}_i - \frac{3\kappa}{4} (\bar{\rho}_i + \bar{p}_i) (\bar{\rho} + \bar{p}) \mathcal{V}_i^2$$

$$+ (\bar{\rho}_i + \bar{p}_i) \delta \Theta \mathcal{V}_i - \frac{\bar{p}_i}{2} (B_\gamma B^\gamma - \phi^2 - 2C\phi) - \frac{\bar{p}_i}{2} (2C_\gamma{}^\nu C_\nu{}^\gamma - C^2).$$

Combining with the gravitational part and discarding the terms proportional to the background equations of motion we obtain:

$$\frac{\delta \mathcal{L}^{(2,s)}}{\sqrt{-g}} = \frac{3\bar{D}^2 \Psi \bar{D}_K^2 \Psi}{\kappa \bar{\Theta}^2} - \sum_i \frac{(\bar{\rho}_i + \bar{p}_i)}{3\kappa(\bar{\rho} + \bar{p})} \Xi_i^2$$

$$+ \sum_i \left[\frac{\bar{c}_{si}^2 \Pi_{\mathcal{V}_i}^2}{2\sqrt{-\bar{g}}^2 (\bar{\rho}_i + \bar{p}_i)} + \frac{9(\bar{\rho}_i + \bar{p}_i)}{2\bar{\Theta}^2} \mathcal{U}_i \bar{D}_K^2 \mathcal{U}_i \right],$$

where

$$\Psi \equiv \psi - \frac{\bar{\Theta}}{3} \delta \sigma^s, \quad \mathcal{U}_i \equiv \psi + \frac{\bar{\Theta}}{3} \mathcal{V}_i, \quad \Xi_i \equiv \delta \Theta - \frac{3\kappa(\bar{\rho} + \bar{p})}{2} \mathcal{V}_i + \frac{3\delta \mathcal{R}}{4\bar{\Theta}}.$$

Constraints Reduction

A close inspection shows that the scalar Lagrangian:

- does not depend on time derivatives of the variables $(\phi, \mathcal{B})$,
- depends quadratically on the time derivatives of $(\psi, \mathcal{E}, \mathcal{U}_i)$.

Using the Faddeev-Jackiv² we perform a Legendre transformation in the time derivatives variables

$$\delta\mathcal{L}^{(2,s)} = \bar{D}_K^2 \Pi_{\mathcal{E}} \mathcal{L}_{\bar{v}} \bar{D}^2 \mathcal{E} + \Pi_{\psi^t} \dot{\psi}^t + \sum_i \bar{D}_K^2 \Pi_{\mathcal{U}_i} \dot{\mathcal{U}}_i - \delta\mathcal{H}_c^{(2,s)},$$

The constrained Hamiltonian $\delta\mathcal{H}_c^{(2,s)}$:

$$\begin{aligned} \delta\mathcal{H}_c^{(2,s)} = & \bar{D}_K^2 \Pi_{\mathcal{E}} \mathcal{L}_{\bar{v}} \bar{D}^2 \mathcal{E} + \Pi_{\psi^t} \dot{\psi}^t + \sum_i \bar{D}_K^2 \Pi_{\mathcal{U}_i} \dot{\mathcal{U}}_i - \sqrt{-\bar{g}} \frac{3\bar{D}^2 \Psi \bar{D}_K^2 \Psi}{\kappa \bar{\Theta}^2} \\ & - \sum_i \left[\frac{\bar{c}_{si}^2 \bar{\Theta}^2 (\bar{D}_K^2 \Pi_{\mathcal{U}_i})^2}{18 \sqrt{-\bar{g}} (\bar{\rho}_i + \bar{p}_i)} - \sqrt{-\bar{g}} \frac{(\bar{\rho}_i + \bar{p}_i)}{3\kappa(\bar{\rho} + \bar{p})} \Xi_i^2 + \sqrt{-\bar{g}} \frac{9(\bar{\rho}_i + \bar{p}_i)}{2\bar{\Theta}^2} \mathcal{U}_i \bar{D}_K^2 \mathcal{U}_i \right]. \end{aligned}$$

²L. Faddeev and R. Jackiw, *Physical Review Letters* 60 (1988), 1692–1694.

Constraints Reduction

The variables in blue are the time derivatives written in terms of the momenta:

$$\begin{aligned}\mathcal{L}_{\bar{v}}[\bar{D}^2\mathcal{E}] &= \bar{D}^2\mathcal{B} - \frac{3\bar{D}^2\psi}{\bar{\Theta}} + \frac{3\bar{D}^2\Psi}{\bar{\Theta}}, \\ \dot{\psi}^t &= \frac{\kappa(\bar{\rho} + \bar{p})}{2}\mathcal{V} - \frac{\bar{\Theta}}{3}\phi - \frac{\bar{D}^2\mathcal{B}}{3} + \frac{\bar{D}_K^2\psi}{\bar{\Theta}} + \frac{\Xi}{3}, \\ \dot{\mathcal{U}}_i &= \frac{\dot{\bar{\Theta}}}{\bar{\Theta}}\mathcal{U}_i + \frac{\bar{c}_{si}^2\bar{\Theta}^2\bar{D}_K^2\Pi_{\mathcal{U}i}}{9\sqrt{-\bar{g}}(\bar{\rho}_i + \bar{p}_i)} + \frac{3\kappa(\bar{\rho} + \bar{p})}{2\bar{\Theta}}\mathcal{U} + \frac{\bar{D}^2\Psi}{\bar{\Theta}} + \frac{\Xi}{3},\end{aligned}$$

where we introduced effective fluid variables

$$(\bar{\rho} + \bar{p})\mathcal{U} = \sum_i (\bar{\rho}_i + \bar{p}_i)\mathcal{U}_i, \quad (\bar{\rho} + \bar{p})\Xi = \sum_i (\bar{\rho}_i + \bar{p}_i)\Xi_i.$$

Therefore, the constraints are simply:

$$\bar{D}_K^2\Pi_{\mathcal{E}} = 0, \quad \Pi_{\psi^t} = 0.$$

Constraints Reduction

Using the expressions for the momenta:

$$\begin{aligned}\bar{D}_K^2 \Pi_{\mathcal{U}i} &\equiv \frac{\partial \delta \mathcal{L}^{(2,s)}}{\partial \dot{\mathcal{U}}_i} = \frac{3\Pi_{\mathcal{V}i}}{\bar{\Theta}}, \\ \bar{D}_K^2 \Pi_{\mathcal{E}} &\equiv \frac{\partial \delta \mathcal{L}^{(2,s)}}{\partial \dot{\mathcal{L}}_{\bar{\nu}}[\bar{D}^2 \mathcal{E}]} = \frac{2\sqrt{-\bar{g}}}{\kappa \bar{\Theta}} \bar{D}_K^2 \Psi - \sum_i \frac{\Pi_{\mathcal{U}i}}{3}, \\ \Pi_{\psi^{\dagger}} &\equiv \frac{\partial \delta \mathcal{L}^{(2,s)}}{\partial \dot{\psi}^{\dagger}} = -2\sqrt{-\bar{g}} \frac{\Xi}{\kappa} - \sum_i \Pi_{\mathcal{U}i},\end{aligned}$$

then

$$\Xi = -\frac{\kappa}{2\sqrt{-\bar{g}}} \sum_i \bar{D}_K^2 \Pi_{\mathcal{U}i}, \quad \Psi = \frac{\kappa \bar{\Theta}}{6\sqrt{-\bar{g}}} \Pi_{\mathcal{U}}, \quad \Pi_{\mathcal{U}} \equiv \sum_i \Pi_{\mathcal{U}i},$$

the equations above also imply

$$\bar{D}_K^2 \Psi = -\frac{\bar{\Theta} \Xi}{3}.$$

Adiabatic and Entropy Perturbations

The calculation above shows that:

- The constraints naturally lead to effective fluids variable.
- The physical degrees of freedom are $(\mathcal{U}_i, \Pi_{\mathcal{U}_i})$.

We incorporate both notions writing the Lagrangian in terms of the effective fluid (curvature perturbation) variable:

$$\zeta \equiv \mathcal{U} - \frac{\bar{K}\bar{\Theta}\Pi_{\mathcal{U}}}{3\sqrt{-\bar{g}}(\bar{\rho} + \bar{p})},$$

and the gauge invariant energy density contrast

$$\delta_{\psi i} = 3\mathcal{U}_i - \frac{\bar{\Theta}\bar{D}_K^2\Pi_{\mathcal{U}_i}}{3\sqrt{-\bar{g}}(\bar{\rho}_i + \bar{p}_i)} = \frac{\delta\rho_i}{\bar{\rho}_i + \bar{p}_i} + 3\psi.$$

Naturally we define

$$(\bar{\rho} + \bar{p})\delta_{\psi} = \sum_i (\bar{\rho}_i + \bar{p}_i)\delta_{\psi i}, \quad \tilde{\delta}_{\psi i} = \delta_{\psi i} - \delta_{\psi}, \quad \tilde{\mathcal{U}}_i \equiv \mathcal{U}_i - \mathcal{U}.$$

Multiple Fluids Hamiltonian

Splitting the Lagrangian between adiabatic and entropy modes

$$\delta\mathcal{L}^{(2,s)} = \delta\mathcal{L}_a^{(2,s)} + \delta\mathcal{L}_s^{(2,s)}.$$

we have

$$\delta\mathcal{L}_a^{(2,s)} = \Pi_\zeta \dot{\zeta} - \delta\mathcal{H}_a^{(2,s)},$$

$$\delta\mathcal{L}_s^{(2,s)} = \sum_i \frac{3\sqrt{-\bar{g}}(\bar{\rho}_i + \bar{p}_i)\tilde{\mathcal{U}}_i}{\bar{\Theta}} \dot{\delta}_{\psi i} - \delta\mathcal{H}_s^{(2,s)},$$

$$\delta\mathcal{H}_a^{(2,s)} = \frac{\bar{c}_s^2 \bar{\Theta}^2 \Pi_\zeta \bar{D}^2 \bar{D}_K^{-2} \Pi_\zeta}{18\sqrt{-\bar{g}}(\bar{\rho} + \bar{p})} - \sqrt{-\bar{g}} \frac{9(\bar{\rho} + \bar{p})}{2\bar{\Theta}^2} \zeta \bar{D}_K^2 \zeta,$$

$$\delta\mathcal{H}_s^{(2,s)} = \sum_i \left[\frac{\sqrt{-\bar{g}} \bar{c}_{si}^2 (\bar{\rho}_i + \bar{p}_i) \tilde{\delta}_{\psi i}^2}{2} - \sqrt{-\bar{g}} \frac{9(\bar{\rho}_i + \bar{p}_i)}{2\bar{\Theta}^2} \tilde{\mathcal{U}}_i \bar{D}^2 \tilde{\mathcal{U}}_i - \frac{\bar{\Theta} \Pi_\zeta}{3(\bar{\rho} + \bar{p})} \bar{c}_{si}^2 (\bar{\rho}_i + \bar{p}_i) \tilde{\delta}_{\psi i} \right].$$

where we defined the momentum conjugated to ζ as $\Pi_\zeta \equiv \bar{D}_K^2 \Pi_{\mathcal{U}}$.

Two Fluids Lagrangian

For two fluids the perturbations $\tilde{\delta}_{\psi i}$ and $\tilde{\mathcal{U}}_i$ can be written in terms of

$$\Pi_Q \equiv \delta_{\psi 2} - \delta_{\psi 1}, \quad \text{and}, \quad Q = \frac{3\sqrt{-\bar{g}}\varpi}{\bar{\Theta}}(\mathcal{U}_1 - \mathcal{U}_2),$$

where the “reduced variable” is $\varpi \equiv (\bar{\rho}_1 + \bar{p}_1)(\bar{\rho}_2 + \bar{p}_2)/(\bar{\rho} + \bar{p})$. The Lagrangian/Hamiltonian then reads

$$\delta\mathcal{L}^{(2,s)} = \Pi_\zeta \zeta' + \Pi_Q Q' - \left[\frac{\Pi_\zeta^2}{2m_\zeta} + \frac{\Pi_Q^2}{2m_S} + \frac{m_\zeta \nu_\zeta^2 \zeta^2}{2} + \frac{m_S \nu_S^2 Q^2}{2} + \frac{\bar{c}_n^2}{\bar{c}_s^2 \bar{c}_m^2} \frac{\Pi_\zeta \Pi_Q}{m_\zeta m_S N H \Delta_K} \right],$$

where we changed the time derivative introducing the lapse function N ,

$$m_\zeta \equiv \frac{a^3(\bar{\rho} + \bar{p})}{N \bar{c}_s^2 \Delta_K H^2}, \quad \nu_\zeta^2 \equiv -\frac{N^2 \bar{c}_s^2 \hat{D}^2}{a^2},$$

$$m_S \equiv \frac{1}{N a^3 \bar{c}_m^2 \varpi}, \quad \nu_S^2 \equiv -\frac{N^2 \bar{c}_m^2 \hat{D}^2}{a^2}.$$

Symplectic structure

To perform the quantization we need the following structure:

- The phase space vector $\chi_a = (\varphi_1, \dots, \varphi_n, \Pi_{\varphi_1}, \dots, \Pi_{\varphi_n})$.
- The Hamiltonian

$$\mathcal{H}(\chi) = \frac{\chi_a \mathcal{H}^{ab} \chi_b}{2},$$

with $\mathcal{H}^{ab}$ being symmetric.

- The symplectic forms

$$\mathbb{S}_{ab} \doteq i \begin{pmatrix} 0 & \mathbf{1}_{n \times n} \\ -\mathbf{1}_{n \times n} & 0 \end{pmatrix}, \quad \mathbb{S}^{ab} \doteq i \begin{pmatrix} 0 & \mathbf{1}_{n \times n} \\ -\mathbf{1}_{n \times n} & 0 \end{pmatrix},$$

- The solutions thus satisfy $i\mathcal{L}_{\bar{v}}\chi_a = \mathbb{S}_{ab}\mathcal{H}^{bc}\chi_c$.
- The product of two solutions χ and ϖ ,

$$\mathbb{S}(\chi, \varpi) = \int_{\Sigma} d^3x \chi_a \varpi_b \mathbb{S}^{ab},$$

is conserved, i.e., $i\mathcal{L}_{\bar{v}}(\chi, \varpi) = 0$.

Complex Phase Space

We complexify the phase space and define the product

$$(\chi, \varpi) \equiv \mathbb{S}(\chi^*, \varpi).$$

To probe the phase space we introduce the Laplacian eigenfunctions

$$\bar{D}^2 \mathcal{Y}_q = -\lambda_q^2 \mathcal{Y}_q, \quad \int_{\Sigma} d^3x \mathcal{Y}_{q_1} \mathcal{Y}_{q_2} = \delta^3(q_1 - q_2).$$

We write the vector in the phase space for each mode as,

$$\begin{aligned} U_{q,a} &\equiv T_a \mathcal{Y}_q, \\ T_a &= (\varphi_{q,1}, \dots, \varphi_{q,n}, \Pi_{\varphi_{q,1}}, \dots, \Pi_{\varphi_{q,n}}), \end{aligned}$$

where the functions T_a depend only on time $\bar{D}_\mu T_a = 0$ and q .

The product of two solution will be given by

$$(U_{q_1}, V_{q_2}) = T_a^*(q_1) W_b(q_2) \mathbb{S}^{ab} \delta^3(q_1 - q_2),$$

where $V_q = W_a \mathcal{Y}_q$.

Finite Dimensional Phase Space

Now our problem is reduced to the product

$$\mathbf{T} \cdot \mathbf{W} = \mathbf{T}_a^* \mathbb{S}^{ab} \mathbf{W}_b,$$

defined in a finite $2n$ -dimensional phase space for each mode q .
If for a given vector $\mathbf{T}_a$ its norm is positive

$$\mathbf{T} \cdot \mathbf{T} = \mathbf{T}_a^* \mathbf{T}_b \mathbb{S}^{ab} > 0,$$

the vector $\mathbf{T}_a^*$ will have the norm

$$\mathbf{T}^* \cdot \mathbf{T}^* = \mathbf{T}_a \mathbf{T}_b^* \mathbb{S}^{ab} = -\mathbf{T} \cdot \mathbf{T} < 0.$$

Thus we build normalized n -dimensional basis such that $e^i \cdot e^j = \delta^{ij}$,
consequently

$$e^{i*} \cdot e^{j*} = -\delta^{ij}, \quad e^{i*} \cdot e^j = e^i_a \mathbb{S}^{ab} e^j_b = 0.$$

Canonical Quantization

The Poisson bracket structure of our problem is given by

$$\{F_1, F_2\} = -i \int_{\Sigma} d^3x \frac{\delta F_1}{\delta \chi_a(x)} \mathbb{S}_{ab} \frac{\delta F_2}{\delta \chi_b(x)},$$

where F_1 and F_2 are two field functionals. Using the definitions above is easy to see that

$$\{\chi_a(x), \chi_b(x')\} = -i \mathbb{S}_{ab} \delta^3(x - x'),$$

where we have, for example, $\{\varphi_1, \Pi_{\varphi_1}\} = \delta^3(x - x')$, as expected. By the canonical quantization rules we promote the fields to Hermitian operators. The fields operators then satisfy the commutation relations,

$$[\hat{\chi}_a(x), \hat{\chi}_b(x')] = \mathbb{S}_{ab} \delta^3(x - x').$$

Creation and Annihilation Operators

First of all, we extend the product to the operators (and classical fields)

$$(\chi, \varpi) = \mathbb{S}(\chi^\dagger, \varpi).$$

Having the following properties:

- $[(\varpi, \hat{\chi}), (\vartheta, \hat{\chi})] = (\vartheta, \varpi^*),$
- $\left[(\varpi, \hat{\chi}), (\vartheta, \hat{\chi})^\dagger\right] = (\varpi, \vartheta),$
- $\left[(\varpi, \hat{\chi})^\dagger, (\vartheta, \hat{\chi})^\dagger\right] = (\vartheta^*, \varpi).$

Given a orthonormal basis $U_{q,a}^i$, the annihilation operators associated with this basis is

$$a_q^i \equiv (U_q^i, \hat{\chi}).$$

It follows directly from the definition and the properties above that

$$\left[a_q^i, a_{q'}^{j\dagger}\right] = \delta^{ij}\delta^3(q - q'), \quad \left[a_q^i, a_{q'}^j\right] = 0 = \left[a_q^{i\dagger}, a_{q'}^{j\dagger}\right].$$

Representation Choice

- Each choice of basis $U_{q,a}^i$ produces a different representation.
- The representation choice reduces to find a basis in the finite dimensional space of T^i .

Defining a vacuum at the instant t_1 , $T_a^i(t_1) = t_a^i$, at a time t_2 we have

$$T_a^i(t_2) = \alpha_j^i(t_2) t_a^j + \beta_j^i(t_2) t_a^{j*},$$

where the functions

$$\alpha_j^i(t) \equiv T_a^i(t) t_j^{a*}, \quad \beta_j^i(t) \equiv T_a^i(t) t_j^a,$$

satisfy $\alpha_j^i(t_1) = \delta_j^i$ and $\beta_j^i(t_1) = 0$.

Then, the annihilation and creation operators at t_2 can be written in terms of the same operators at t_1 as

$$\begin{aligned} a_q^i(t_2) &= \alpha_j^{i*}(t_2) a_q^j(t_1) - \beta_j^{i*}(t_2) a_q^{j\dagger}(t_1), \\ a_q^{i\dagger}(t_2) &= \alpha_j^i(t_2) a_q^{j\dagger}(t_1) - \beta_j^i(t_2) a_q^j(t_1). \end{aligned}$$

Representation Choice

- The number operator $N_q^i(t) \equiv a_q^{i\dagger}(t)a_q^i(t)$ applied at $|0_{t_1}\rangle$ measures

$$\langle 0_{t_1} | N_q^i(t) | 0_{t_1} \rangle = \delta^3(0) \int d^3q \sum_j |\beta_j^i(t)|^2,$$

where the $\delta^3(0)$ is volume of the spatial section.

- If the integral converges we have defined a unitary evolution.
- The time evolution of each matrix are

$$i\dot{\alpha}_j^i = M_k^i \alpha_j^k - N_k^i \beta_j^{k*},$$

$$i\dot{\beta}_j^i = M_k^i \beta_j^k - N_k^i \alpha_j^{k*},$$

where we defined the following matrices

$$M_k^i \equiv T_a^i \mathcal{H}^{ab} T_{kb}^*, \quad N_k^i \equiv T_a^i \mathcal{H}^{ab} T_{kb}.$$

- Note that the matrix N_j^k control the mixing between α_j^i and β_j^i , for instance, **if it is null then there is no particle creation**.
- However, in general $\mathcal{H}_a^b$ will depend on time, thus, even if N_j^k is null initially at t_1 nothing guarantees that it will remain null.

Hamiltonian Eigenvectors

- The matrix $\mathcal{H}_a{}^b \equiv \mathbb{S}_{ac} \mathcal{H}^{cb}$ works as an operator in the space of the solutions.
- This operator is Hermitian, given the product of the two vectors V_a and $\mathcal{H}_a{}^b U_b$,

$$V \cdot (\mathcal{H}U) = V_a^* \mathbb{S}^{ab} (\mathcal{H}_b{}^c U_c) = (\mathcal{H}_c{}^a V_a)^* \mathbb{S}^{cd} U_d = (\mathcal{H}V) \cdot U.$$

- Finally, if we assume that $\mathbf{t}_a^i$ are the eigenvectors of $\mathcal{H}_a{}^b(t_1)$, the condition of the time independent vacuum is simply

$$N^{ij}(t_1) = \mathbf{t}_a^i \mathbb{S}^{ab} \mathcal{H}_b{}^c \mathbf{t}_c^j \big|_{t_1} = \nu_j^j \mathbf{t}_a^i \mathbb{S}^{ab} \mathbf{t}_b^k \big|_{t_1} = 0,$$

where $\mathcal{H}_a{}^b \mathbf{t}_b^i = \nu_j^i \mathbf{t}_a^j$ and ν_j^j is a diagonal real matrix containing the eigenvalues.

- This gives zero particle creation in first order expansion of $N^i_k(t)$.
- This choice **does not** guarantee that higher order terms will be zero or provides a convergent β^i_j .

Diagonal $\mathcal{H}^{ab}$

For a diagonal Hamiltonian

$$\mathcal{H}^{ab} \doteq \text{diag}(m_1\nu_1^2, \dots, m_n\nu_n^2, 1/m_1, \dots, 1/m_n).$$

The eigenvectors are given by

$$V^i_a \doteq \left(\frac{1}{\sqrt{2m_i\nu_i}}, -i\sqrt{\frac{m_i\nu_i}{2}} \right).$$

- The canonical transformation in each field in the form

$$q_i \rightarrow \sqrt{\frac{m_i}{M_i}} q_i, \quad p_i \rightarrow M_i \left(\sqrt{\frac{m_i}{M_i}} \right)' q_i + \sqrt{\frac{M_i}{m_i}} p_i.$$

- It changes the masses arbitrarily $m_i \rightarrow M_i$.
- The frequency, however, change as

$$\nu_i^2 \rightarrow W_i^2 = \nu_i^2 - \frac{(\sqrt{m_i})''}{\sqrt{m_i}} + \frac{(\sqrt{M_i})''}{\sqrt{M_i}}.$$

Diagonal $\mathcal{H}^{ab}$

In the new representation the eigenvectors are simply

$$V^i_a \doteq \left(\frac{1}{\sqrt{2M_i W_i}}, -i\sqrt{\frac{M_i W_i}{2}} \right).$$

Then we can make two main choices for

$$M_i W_i = \sqrt{M_i^2 \nu_i^2 - M_i^2 \left(\frac{(\sqrt{m_i})''}{\sqrt{m_i}} - \frac{(\sqrt{M_i})''}{\sqrt{M_i}} \right)}.$$

- Constant $M_i^2 \nu_i^2$, defines algebraically M_i for each field and it is possible to show that leads to a unitary evolution with particle creation.
- Constant $M_i W_i$, leads to a differential equation definition of W_i , i.e.,

$$W_i^2 + \frac{1}{2} \frac{W_i''}{W_i} - \frac{3}{4} \frac{W_i'^2}{W_i^2} = \nu_i^2 - \frac{(\sqrt{m_i})''}{\sqrt{m_i}}.$$

In this representation the eigenvectors are constant, and is possible to show that it leads to unitary evolution with **zero** particle creation.

Adiabatic Vacuum

- The last option leads naturally to the adiabatic vacuum, the differential equation can be approximated by a asymptotic series starting with

$$W_i^2 \approx \nu_i^2 - \frac{(\sqrt{m_i})''}{\sqrt{m_i}}.$$

- This is exactly what one would obtain from the WKB approximation of the field, for **any** field representation.
- This can be extended to non-diagonal Hamiltonians, and is there is a limit where the Hamiltonian is diagonal, we can show that there is also a representation with zero particles production.

To perform the quantization we take the last steps and make the canonical transformation:

$$A = \sqrt{\frac{m_\zeta}{m_A}} \zeta, \quad P_A = \sqrt{\frac{m_A}{m_\zeta}} \Pi_\zeta + m_A \sqrt{\frac{m_\zeta'}{m_A}} \zeta,$$

$$B = \sqrt{\frac{m_S}{m_B}} Q, \quad P_B = \sqrt{\frac{m_B}{m_S}} \Pi_Q + m_B \sqrt{\frac{m_S'}{m_B}} Q,$$

With the new frequencies satisfying

$$\nu_A^2 + \frac{1}{2} \frac{\nu_A''}{\nu_A} - \frac{3}{4} \frac{\nu_A'^2}{\nu_A^2} = \nu_\zeta^2 - \frac{(\sqrt{m_\zeta})''}{\sqrt{m_\zeta}},$$

$$\nu_B^2 + \frac{1}{2} \frac{\nu_B''}{\nu_B} - \frac{3}{4} \frac{\nu_B'^2}{\nu_B^2} = \nu_S^2 - \frac{(\sqrt{m_S})''}{\sqrt{m_S}}.$$

The Hamiltonian in these variables is

$$\delta \mathcal{L}^{(2,s)} = P_A A' + P_B B' - \left[\frac{\nu_A P_A^2}{2 \tilde{\lambda}_q} + \frac{\nu_B P_B^2}{2 \tilde{\lambda}_q} \right. \\ \left. + \frac{\tilde{\lambda}_q \nu_A A^2}{2} + \frac{\tilde{\lambda}_q \nu_B B^2}{2} + y(P_A - L_A A)(P_B - L_B B) \right].$$

Integral Solution

Writing the fields in terms of

$$R_A^\pm = \sqrt{\tilde{\lambda}_q} A \pm i \frac{P_A}{\sqrt{\tilde{\lambda}_q}} \quad \text{and} \quad R_B^\pm = \sqrt{\tilde{\lambda}_q} B \pm i \frac{P_B}{\sqrt{\tilde{\lambda}_q}}.$$

The equations for these variables can be readily integrated, resulting in a set of integral equations,

$$R_A^\pm = e^{\mp i \int d\tau \nu_A} \left[R_{A0}^\pm + \int d\tau y \left(\sqrt{\tilde{\lambda}_q} \pm i \frac{L_A}{\sqrt{\tilde{\lambda}_q}} \right) (P_B - L_B B) e^{\pm i \int d\tau \nu_A} \right],$$

$$R_B^\pm = e^{\mp i \int d\tau \nu_B} \left[R_{B0}^\pm + \int d\tau y \left(\sqrt{\tilde{\lambda}_q} \pm i \frac{L_B}{\sqrt{\tilde{\lambda}_q}} \right) (P_A - L_A A) e^{\pm i \int d\tau \nu_B} \right].$$

Adiabatic Approximation

If ν_A and ν_B are large enough, we can approximate the integral solutions by

$$R_A^\pm = e^{\mp i \int d\tau \nu_A} R_{A0}^\pm \pm \frac{1}{i\nu_A} y \left(\sqrt{\tilde{\lambda}_q} \pm i \frac{L_A}{\sqrt{\tilde{\lambda}_q}} \right) (P_B - L_B B),$$

$$R_B^\pm = e^{\mp i \int d\tau \nu_B} R_{B0}^\pm \pm \frac{1}{i\nu_B} y \left(\sqrt{\tilde{\lambda}_q} \pm i \frac{L_B}{\sqrt{\tilde{\lambda}_q}} \right) (P_A - L_A A).$$

- This approximation assumes only that ν_A and ν_B are large.
- Therefore, it can be applied even when the coupling y is large.
- The equations above provide a linear system in terms of $R_A^\pm$ and $R_B^\pm$.

Conclusions

- We obtained the second order Hamiltonian for a multiple fluids system.
- This Hamiltonian was obtained without assuming a dynamics for the background.
- We suggested a well defined procedure to find a vacuum for a multiple components system.
- This in turn leads to a WKB approximation which can be easily applied to a multiple components system.