

Spherically symmetric solutions of Massive gravity in the Decoupling Limit

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based on arXiv:0901.0393 + work in progress



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Outline of the presentation

- Introduction & definitions
- The Goldstone boson formalism and the Decoupling Limit
- Spherically symmetric solutions in the Decoupling Limit
- Conclusion

Introduction

Context & Motivations

- **Modified gravity theories** - a way to get late time acceleration of the Universe and (maybe) solve the Cosmological Constant problem.
➡ **Naive idea: give a mass to the graviton with $m \sim 1/H_0$**
- Pathologies: non-bounded by below Hamiltonian and ghosts [Boulware & Deser '72; Creminelli, Nicolis, Papucci, Trincherini '05; Deffayet & Rombouts '05], singular solutions (?) [Damour et al. '03], but interesting toy model for more realistic theories.
- Massive gravity = basic ingredient for models with extra-dimensions (like DGP):
 - in 5D, the gravity is massless,
 - viewed from the 4D brane, the gravity is mediated by a tower a massive Kaluza-Klein gravitons with no massless mode.

What is Massive gravity?

- The quadratic action for the massive graviton:

f : background metric (often flat)

$H_{\mu\nu}$: spin 2 excitation over f

$$S = \frac{M_P^2}{2} \int d^4x \left(\underbrace{“H \partial^2 H + …”}_{\text{Kinetic term}} - \frac{m^2}{4} \underbrace{[H_{\mu\nu} H^{\mu\nu} - (H^\mu_\mu)^2]}_{\text{Pauli-Fierz Mass term}} \right) + \int d^4x \frac{1}{2} \underbrace{T_{\mu\nu} H^{\mu\nu}}_{\text{Matter coupling}}$$

- Non-linear completion: dynamical metric $g = f + H$

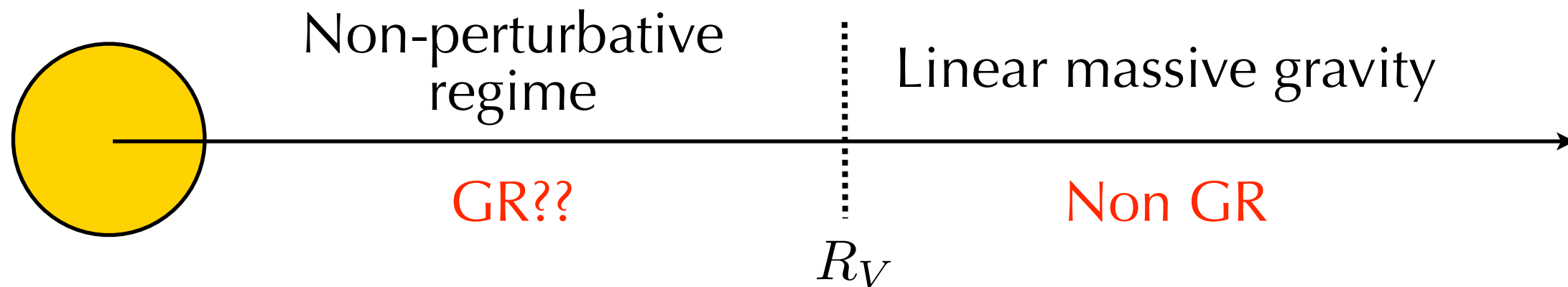
$$S = \frac{M_P^2}{2} \int d^4x \left(\underbrace{\sqrt{-g} R[g]}_{\text{Scalar density}} \right) + S_m[g]$$

- Examples:

$$\mathcal{V}^{(BD)}[\mathbf{g}^{-1}\mathbf{f}] = \sqrt{-f} H_{\mu\nu} H_{\sigma\tau} (f^{\mu\sigma} f^{\nu\tau} - f^{\mu\nu} f^{\sigma\tau}) \quad \text{Boulware and Deser '72}$$

$$\mathcal{V}^{(AGS)}[\mathbf{g}^{-1}\mathbf{f}] = \sqrt{-g} H_{\mu\nu} H_{\sigma\tau} (g^{\mu\sigma} g^{\nu\tau} - g^{\mu\nu} g^{\sigma\tau}) \quad \text{Arkani-Hamed, Georgi and Schwartz '03}$$

Static Spherically Symmetric solutions



- New scale: the Vainshtein radius $R_V = (R_S m^{-4})^{1/5}$ *Vainshtein '72*

➡ Is it possible to find a solution regular everywhere?

- Our approach is to study these questions in a specific limit: the decoupling limit

$$\begin{array}{c|c} M_P \rightarrow \infty & \Lambda \equiv (M_P m^4)^{1/5} \sim \text{const} \\ m \rightarrow 0 & T_{\mu\nu}/M_P \sim \text{const} \end{array}$$

Arkani-Hamed, Georgi and Schwartz '03

➡ Are there regular solutions in the DL?

➡ To what extent does the DL encode the physics of the full system?

The Goldstone Picture & the Decoupling Limit

The Stuckelberg mechanism (I)

Goal: to separate explicitly the various degrees of freedom (tensor, vector, scalar) of a massive field.

Example: *the Proca's field* $\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} - \frac{m^2}{2}A_\mu A^\mu$ ← breaks gauge invariance

- Field redefinition $A_\mu \rightarrow A_\mu - \partial_\mu B$

$$\Rightarrow \mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} - \frac{1}{2}m^2(A_\mu - \partial_\mu B)(A^\mu - \partial^\mu B)$$

New gauge invariance: $A_\mu \rightarrow A_\mu + \partial_\mu \Lambda$
 $B \rightarrow B + \Lambda$

- “Unitary gauge”: $B = 0$
- “Longitudinal gauge”: $\partial_\mu A^\mu = 0$

$$\Rightarrow \mathcal{L} = \overset{2 \text{ DOF}}{-\frac{1}{2}\partial_\mu A_\nu \partial^\mu A^\nu} - \frac{m^2}{2}A_\mu A^\mu - \overset{1 \text{ DOF}}{\frac{m^2}{2}\partial_\mu B \partial^\mu B}$$

The Stuckelberg mechanism (II)

- Massive spin-2 graviton: in the action

$$S = \frac{M_P^2}{2} \int d^4x \left(\sqrt{-g} R[g] - \frac{m^2}{4} \mathcal{V} [\mathbf{g}^{-1} \mathbf{f}] \right) + S_m[g],$$

replace $f_{\mu\nu}(x) \rightarrow f_{\mu\nu}(x) = \partial_\mu X^A(x) \partial_\nu X^B(x) f_{AB}(X(x))$
 $g_{\mu\nu}(x) \rightarrow g_{\mu\nu}(x)$

the action is now invariant under both $X^A \rightarrow X'^A$
 $x^\mu \rightarrow x'^\mu$

- “Unitary gauge”: $X_0^A(x) \equiv \delta_\mu^A x^\mu$
- In non-unitary gauge, introduce the “Goldstone boson” π :

$$X^A(x) = X_0^A(x) + \pi^A(x).$$

The Goldstone boson expansion

- Scalar-vector decomposition: $\pi^A(x) = f^{AB} (A_B + \partial_B \phi)$
- In terms of $h_{\mu\nu} = g_{\mu\nu} - f_{\mu\nu}$, it corresponds to the redefinition:

$$h_{\mu\nu} \rightarrow H_{\mu\nu} = h_{\mu\nu} - \partial_\mu A_\nu - \partial_\nu A_\mu - 2\partial_\mu \partial_\nu \phi \\ - \partial_\mu A_\sigma \partial_\nu A^\sigma - \partial_\mu \partial_\sigma \phi \partial_\nu \partial^\sigma \phi \\ - \partial_\nu A^\sigma \partial_\mu \partial_\sigma \phi - \partial_\mu A^\sigma \partial_\nu \partial_\sigma \phi$$

$$\Rightarrow S = \frac{M_P^2}{8} \int d^4x \left\{ 2h^{\mu\nu} \partial_\mu \partial_\nu h - 2h^{\mu\nu} \partial_\nu \partial_\sigma h_\mu^\sigma + h^{\mu\nu} \square h_{\mu\nu} - h \square h \right. \\ \left. + m^2 [h^2 - h_{\mu\nu} h^{\mu\nu} - F_{\mu\nu} F^{\mu\nu} - 4(h \partial A - h_{\mu\nu} \partial^\mu A^\nu) - 4(h \square \phi - h_{\mu\nu} \partial^\mu \partial^\nu \phi)] \right\} \\ + \int d^4x \frac{1}{2} T_{\mu\nu} h^{\mu\nu}$$

Mixing

Shift: $h_{\mu\nu} = \hat{h}_{\mu\nu} - m^2 \eta_{\mu\nu} \phi$ and gauge fixing ➡ demixing

The Decoupling Limit

- Form of the kinetic term:

$$S \supset \int d^4x \left\{ M_P^2 \hat{h} \square \hat{h} + \dots + M_P^2 m^2 A \square A + \dots + M_P^2 m^4 \phi \square \phi + \dots \right\}$$

- Canonical normalization: $\tilde{h}_{\mu\nu} = M_P \hat{h}_{\mu\nu},$
 $\tilde{A}^\mu = M_P m A^\mu,$
 $\tilde{\phi} = M_P m^2 \phi.$

- Dominant higher order term: $\frac{(\partial^2 \tilde{\phi})^3}{\Lambda^5}$ with $\Lambda = (m^4 M_P)^{1/5}$

- In order to concentrate on these terms, take the **Decoupling Limit**:

$$\begin{aligned} M_P &\rightarrow \infty \\ m &\rightarrow 0 \\ \Lambda &\sim \text{const} \\ T_{\mu\nu}/M_P &\sim \text{const} \end{aligned}$$

Action for $\tilde{\phi}$ in the Decoupling Limit

- The action for the scalar sector:

$$S = \frac{1}{2} \int d^4x \left\{ \frac{3}{2} \tilde{\phi} \square \tilde{\phi} + \frac{1}{\Lambda^5} \left[\alpha (\square \tilde{\phi})^3 + \beta (\square \tilde{\phi} \tilde{\phi}_{,\mu\nu} \tilde{\phi}^{,\mu\nu}) \right] - \frac{1}{M_P} T \tilde{\phi} \right\}$$

- Equation of Motion:

$$\nabla_\mu \left\{ 3\Lambda^5 \nabla^\mu \tilde{\phi} + 3\alpha \nabla^\mu (\square \tilde{\phi})^2 + \beta \nabla^\mu (\tilde{\phi}_{;\delta\gamma})^2 + 2\beta \nabla^\nu (\square \tilde{\phi} \tilde{\phi}_{;\nu}^\mu) \right\} = \frac{\Lambda^5}{M_P} T$$

Can be
integrated!

Spherically Symmetric case:

$$\begin{aligned} 3 \frac{\tilde{\phi}'}{R} + \frac{2}{\Lambda^5} \left\{ 3\alpha \left(-4 \frac{\tilde{\phi}'^2}{R^4} + 2 \frac{\tilde{\phi}' \tilde{\phi}''}{R^3} + 2 \frac{\tilde{\phi}''^2}{R^2} + 2 \frac{\tilde{\phi}' \tilde{\phi}^{(3)}}{R^2} + \frac{\tilde{\phi}'' \tilde{\phi}^{(3)}}{R} \right) + \right. \\ \left. + \beta \left(-6 \frac{\tilde{\phi}'^2}{R^4} + 2 \frac{\tilde{\phi}' \tilde{\phi}''}{R^3} + 4 \frac{\tilde{\phi}''^2}{R^2} + 2 \frac{\tilde{\phi}' \tilde{\phi}^{(3)}}{R^2} + 3 \frac{\tilde{\phi}'' \tilde{\phi}^{(3)}}{R} \right) \right\} \\ = -\frac{1}{R^3} \int_0^R d\tilde{R} \tilde{\rho}(\tilde{R}) \tilde{R}^2 \end{aligned}$$

The Vainshtein mechanism

- Schematic equation of motion for $\tilde{\phi}$:

$$\cancel{\Box \tilde{\phi}} + \frac{1}{\Lambda^5} \Box \left(\Box \tilde{\phi} \right)^2 = \frac{T}{M_P}$$

Strong coupling

- Einstein equations:

$$\tilde{G}_{\mu\nu} + \cancel{\Box \tilde{\phi}} = \frac{T}{M_P}$$

GR!

Spherically symmetric solutions in the Decoupling Limit

The EoM in the Decoupling Limit

- New rescaled variables: $\xi \equiv \frac{R}{R_V}$, $w \equiv -\Lambda^5 R_V^2 \times 2 \frac{\tilde{\phi}'}{R}$

- Equations of Motion:

Linear term

$$3 \frac{\tilde{\phi}'}{R} + \frac{2}{\Lambda^5} \left\{ 3\alpha \left(-4 \frac{\tilde{\phi}'^2}{R^4} + 2 \frac{\tilde{\phi}' \tilde{\phi}''}{R^3} + 2 \frac{\tilde{\phi}''^2}{R^2} + 2 \frac{\tilde{\phi}' \tilde{\phi}^{(3)}}{R^2} + \frac{\tilde{\phi}'' \tilde{\phi}^{(3)}}{R} \right) + \right. \\ \left. + \beta \left(-6 \frac{\tilde{\phi}'^2}{R^4} + 2 \frac{\tilde{\phi}' \tilde{\phi}''}{R^3} + 4 \frac{\tilde{\phi}''^2}{R^2} + 2 \frac{\tilde{\phi}' \tilde{\phi}^{(3)}}{R^2} + 3 \frac{\tilde{\phi}'' \tilde{\phi}^{(3)}}{R} \right) \right\} \\ = -\frac{1}{R^3} \int_0^R d\tilde{R} \tilde{\rho}(\tilde{R}) \tilde{R}^2$$

Source term

Q=quadratic, second order
differential operator

➡ Equation for w:

$$2 Q(w) + \frac{3}{2} w = \frac{1}{\xi^3}$$

$$2Q(w) + \frac{3}{2} w = \frac{1}{\xi^3}$$

close to source

Vainshtein solution,

$$\cancel{2Q(w) + \frac{3}{2} w = \frac{1}{\xi^3}}$$

$$w \propto \frac{1}{\sqrt{\xi}}, \quad (\nu = -\lambda)$$

far from source

perturbative regime,

$$\cancel{2Q(w) + \frac{3}{2} w = \frac{1}{\xi^3}}$$

$$w \rightarrow \frac{2}{3\xi^3}, \quad (\nu = -2\lambda)$$

Another solution!

$$\cancel{2Q(w) + \frac{3}{2} w = \frac{1}{\xi^3}}$$

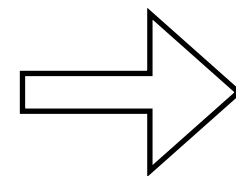
$$Q(w) = 0 \quad \Rightarrow \quad w \propto \xi^{p_{1,2}}, \quad (\nu = -\lambda)$$

Cauchy problem and initial conditions

Is it enough to know the asymptotic behavior at infinity?

NOTE: EOM is singular at $\xi = \infty$

$$2Q(w) + \frac{3}{2} w = \frac{1}{\xi^3}$$
$$w(\infty) \rightarrow \frac{2}{3\xi^3}$$



existence, uniqueness ?

YES (Toy example 1)

$$\begin{cases} w'' + w = \frac{1}{\xi} \\ w(\infty) \rightarrow \frac{1}{\xi} \end{cases}$$

$$w = \frac{1}{\xi} - \frac{2}{\xi^3} + \dots$$
$$+ \cancel{C_1 \cos \xi} + \cancel{C_2 \cos \xi}$$

NO (Toy example 2)

$$\begin{cases} -w'' + w = \frac{1}{\xi} \\ w(\infty) \rightarrow \frac{1}{\xi} \end{cases}$$

$$w = \frac{1}{\xi} - \frac{2}{\xi^3} + \dots$$
$$+ \cancel{C_1 \exp(-\xi)} + \cancel{C_2 \exp \xi}$$

Example 1: the BD potential (I)

$$\mathcal{V}^{(BD)}[\mathbf{g}^{-1}\mathbf{f}] = \sqrt{-f} h_{\mu\nu} h_{\sigma\tau} (f^{\mu\sigma} f^{\nu\tau} - f^{\mu\nu} f^{\sigma\tau}) \quad \text{Boulware, Deser '72}$$

$$2 \left(\frac{\dot{w}^2}{4} + \frac{w\ddot{w}}{2} + 2 \frac{w\dot{w}}{\xi} \right) + \frac{3}{2}w = \frac{1}{\xi^3}$$

- Asymptotic behavior at infinity: **linear theory**

$$\cancel{2 \left(\frac{\dot{w}^2}{4} + \frac{w\ddot{w}}{2} + 2 \frac{w\dot{w}}{\xi} \right) + \frac{3}{2}w = \frac{1}{\xi^3}} \quad \Rightarrow w(\xi) \sim w_{\infty}(\xi) \equiv \frac{2}{3\xi^3}$$

- Linearization around $w_{\infty}(\xi)$

$$w = w_{\infty} + \delta w \Rightarrow \delta w'' + \frac{\delta w'}{\xi} + \frac{9}{4}\xi^3 \delta w = -\frac{3}{\xi^5}$$

~~two oscillatory
modes~~

➡ **unique solution**

Example 1: the BD potential (II)

$$2 \left(\frac{\dot{w}^2}{4} + \frac{w\ddot{w}}{2} + 2\frac{w\dot{w}}{\xi} \right) + \frac{3}{2}w = \frac{1}{\xi^3}$$

- Small distance behavior: **no Vainshtein scaling**

$$2 \left(\frac{\dot{w}^2}{4} + \frac{w\ddot{w}}{2} + 2\frac{w\dot{w}}{\xi} \right) + \cancel{\frac{3}{2}w} = \frac{1}{\xi^3} \Rightarrow \text{try } w(\xi) \sim \frac{A}{\sqrt{\xi}} \rightarrow \text{Imaginary solution!}$$

- Another scaling is possible:

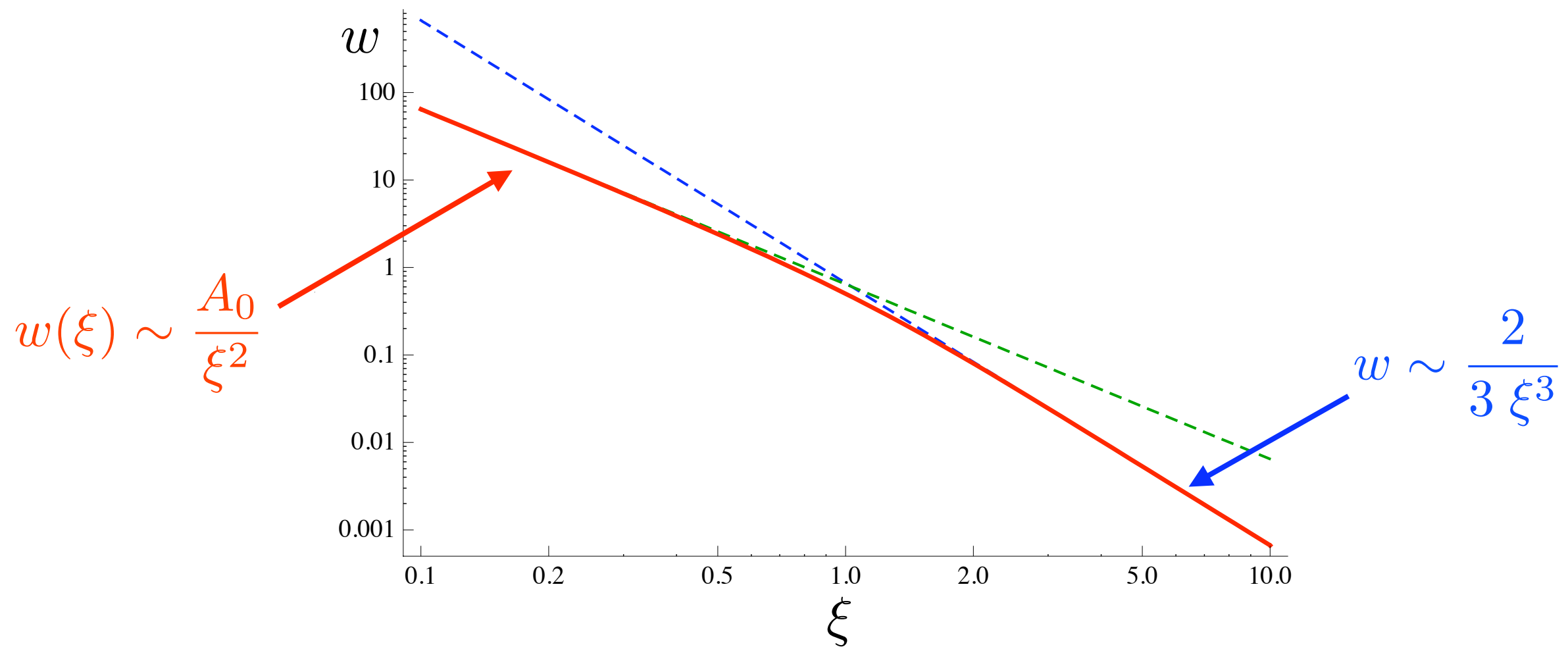
$$2 \left(\frac{\dot{w}^2}{4} + \frac{w\ddot{w}}{2} + 2\frac{w\dot{w}}{\xi} \right) + \cancel{\frac{3}{2}w} = \cancel{\frac{1}{\xi^3}} \Leftrightarrow Q(w) = 0 \Leftrightarrow w(\xi) \sim \frac{A}{\xi^2}$$

2 free constants

$$w(\xi) = \frac{A_0}{\xi^2} + \frac{3A_0 B_0 + \ln \xi}{3A_0} \xi - \frac{3}{8} \xi^2 + \dots$$

Example 1: the BD potential (III)

$$2 \left(\frac{\dot{w}^2}{4} + \frac{w\ddot{w}}{2} + 2\frac{w\dot{w}}{\xi} \right) + \frac{3}{2}w = \frac{1}{\xi^3}$$



➡ Unique regular solution (A_0 and B_0 fixed)

Example 2: the AGS potential (I)

$$\mathcal{V}^{(AGS)}[\mathbf{g}^{-1}\mathbf{f}] = \sqrt{-g} h_{\mu\nu} h_{\sigma\tau} (g^{\mu\sigma} g^{\nu\tau} - g^{\mu\nu} g^{\sigma\tau}) \quad \text{Arkani-Hamed, Georgi, Schwartz '72}$$

$$-2 \left(\frac{\dot{w}^2}{4} + \frac{w\ddot{w}}{2} + 2\frac{w\dot{w}}{\xi} \right) + \frac{3}{2}w = \frac{1}{\xi^3}$$

- Asymptotic behavior at infinity: **linear theory**

~~$$2 \left(\frac{\dot{w}^2}{4} + \frac{w\ddot{w}}{2} + 2\frac{w\dot{w}}{\xi} \right) + \frac{3}{2}w = \frac{1}{\xi^3} \quad \Rightarrow w(\xi) \sim w_{\infty}(\xi) \equiv \frac{2}{3\xi^3}$$~~

- Linearization around $w_{\infty}(\xi)$

$$w = w_{\infty} + \delta w \Rightarrow \delta w'' + \frac{\delta w'}{\xi} - \frac{9}{4}\xi^3 \delta w = -\frac{3}{\xi^5}$$

1 decreasing mode
+ 1 ~~exploding mode~~

➡ solutions isn't unique?

NB: the exploding mode makes the numerical integration tricky!

Example 2: the AGS potential (II)

$$-2 \left(\frac{\dot{w}^2}{4} + \frac{w\ddot{w}}{2} + 2\frac{w\dot{w}}{\xi} \right) + \frac{3}{2}w = \frac{1}{\xi^3}$$

- Small distance behavior: **there is Vainshtein solution**

free constant

$$2 \left(\frac{\dot{w}^2}{4} + \frac{w\ddot{w}}{2} + 2\frac{w\dot{w}}{\xi} \right) + \cancel{\frac{3}{2}w} = \frac{1}{\xi^3} \Rightarrow w(\xi) = \sqrt{\frac{8}{9\xi}}$$

$$w(\xi) = \sqrt{\frac{8}{9\xi}} + \underbrace{B_0}_{\text{free constant}} \xi^{-\frac{5}{4} + \frac{3\sqrt{5}}{4}} - \frac{3(-5 + \sqrt{5})}{8\sqrt{2}(-4 + \sqrt{5})} B_0^2 \xi^{-2 + \frac{3\sqrt{5}}{2}} + \frac{6\xi^2}{31} + \dots$$

- Another scaling is possible:

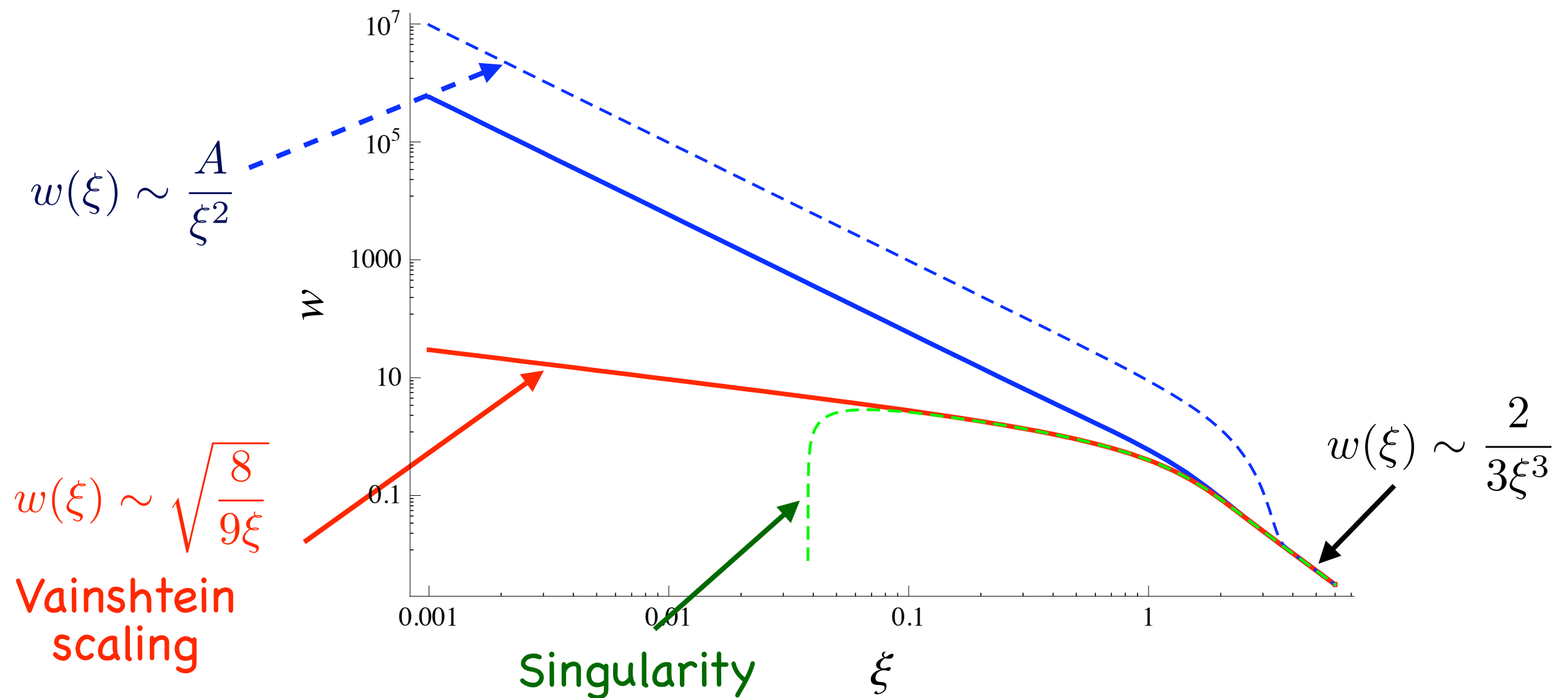
2 free constants

$$2 \left(\frac{\dot{w}^2}{4} + \frac{w\ddot{w}}{2} + 2\frac{w\dot{w}}{\xi} \right) + \cancel{\frac{3}{2}w} = \cancel{\frac{1}{\xi^3}} \Leftrightarrow Q(w) = 0 \Leftrightarrow w(\xi) \sim \frac{A}{\xi^2}$$

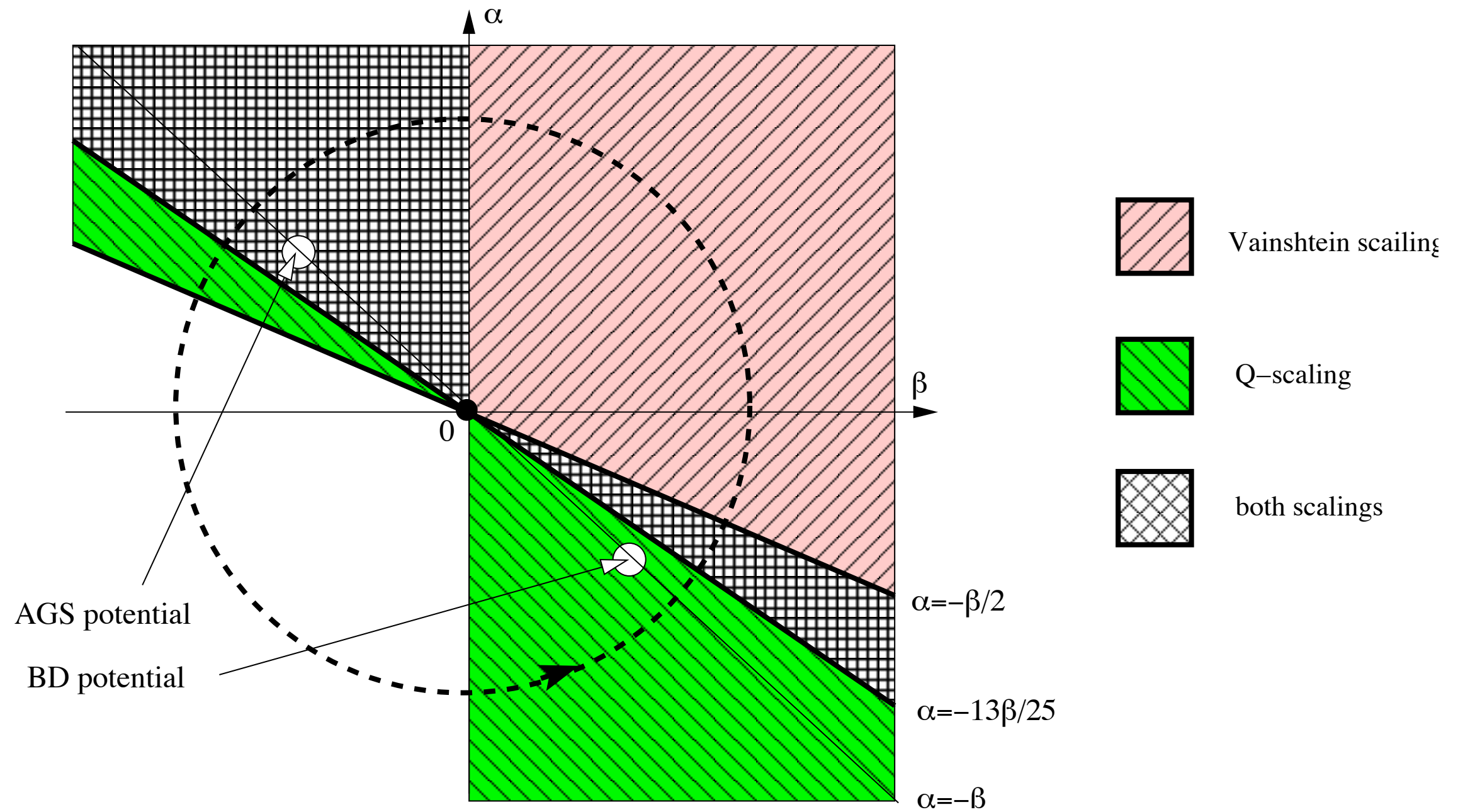
$$w(\xi) = \underbrace{A_0}_{\text{free constant}} \xi^{-2} + \frac{3A_0 \underbrace{B_0}_{\text{free constant}} - \ln \xi}{3A_0} \xi + \frac{3}{8} \xi^2 + \dots$$

Example 2: the AGS potential (III)

NUMERICS



General case of α and β



Conclusion

Conclusion

- The decoupling limit, which focuses on the relevant higher order terms, is an efficient way to classify the different models, and study them.
- We were able to find regular solutions, with two different scalings: Vainshtein and Q-scalings.
- The Vainshtein mechanism works in the DL.
- What remains to be done: see to what extent such solutions can be generalized to the full system (work under progress).