

# Massive Gravity, Lagrangian Media and Thermodynamics of Perfect Fluids

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- Lagrangian of 4 scalar fields  $\mathcal{L}(\partial\Phi^{A=0,1,2,3})$

Stuckelberg fields for spont. broken space-time symm.;

$\phi^{a=1,2,3}$  comoving coordinates of the continuous medium;

$\phi^0$  internal time of the medium

- Internal Symmetries: shift,  $SO(3)_\phi$ , Non Linear extensions
  - Symmetries and Media (Supersolids, solids, superfluids, fluids)
  - Media Lagrangian  $\subset$  Massive Gravity
- Lagrangian formulation of Perfect Fluids
    - Perfect Fluid Lagrangian  $\leftrightarrow$  Thermodynamics
    - Some applications: relativistic Bose-Fermi gas, van der Waals gas, barotropic fluids
    - Thermodynamical stability versus Dynamical stability

# Building the Lagrangian

$$\phi^{A=0,1,2,3} \quad \Rightarrow \quad \partial\phi^A$$

*shift sym.*:  $\phi^a \rightarrow \phi^A + \partial c^A$

$$\underbrace{\Rightarrow}_{\text{Lorentz Scalar}} C^{AB} = \partial_\mu \phi^A g^{\mu\nu} \partial_\nu \phi^B \quad \rightarrow \quad \underbrace{C^{00} \equiv \mathbf{S}}_{\text{Scalar}}, \quad \underbrace{C^{0a} \equiv \mathbf{V}}_{\text{Vector}}, \quad \underbrace{C^{ab} \equiv \mathbf{T}}_{\text{Tensor}} :$$

$SO(3)_\phi$

$$\mathbf{S}, \quad \text{Tr}[\mathbf{T}^{n=1,2,3}], \quad \mathbf{V} \cdot \mathbf{T}^{n=0,1,2,3} \cdot \mathbf{V}$$

$$\underbrace{\Rightarrow}_{\text{Lorentz Vector, } SO(3)_\phi \text{ Scalar}} u^\mu \sim \epsilon^{\mu\alpha\beta\gamma} e_{abc} \partial_\alpha \phi^a \partial_\beta \phi^b \partial_\gamma \phi^c \quad \rightarrow \quad \underbrace{u^\mu \partial_\mu \phi^0}_{\text{Lorentz Scalar}}$$

9 operators

$$A, B = 0, 1, 2, 3, \quad a, b, c = 1, 2, 3$$

# Lagrangian of 4 scalar + shift sym.

- At lowest order in derivatives:  $C^{AB} = \partial_\mu \Phi^A g^{\mu\nu} \partial_\nu \Phi^A$

$$S = \int d^4x \sqrt{-g} (M_{pl}^2 R + U(C^{AB}))$$

Eqs of Motion:

$$\underbrace{\left( U_{AB} g^{\mu\nu} + \left( U_{AB, CB} + U_{AD, BC} \nabla^\mu \Phi^C \nabla^\nu \Phi^D \right) \right)}_{\text{Rank}[\%] = \# \text{ DoF}} \nabla_\mu \nabla_\nu \Phi^B = 0 \quad \leftrightarrow \quad \nabla^\nu T_{\nu\mu} = 0$$

$$U_{AB} = \partial U / \partial C^{AB}, \quad U_{AB, CD} = \partial^2 U / (\partial C^{AB} \partial C^{CD})$$

# Lagrangian of 4 scalar + shift + $SO(3)_\phi$ symm.

global internal spatial  $SO(3)_\phi$  symmetry

$$\phi^0 \rightarrow \phi^0, \quad \phi^a \rightarrow R^a_b \phi^b \quad R R^T = I, \quad a = 1, 2, 3$$

$SO(3)_\phi$  Tensors :  $C^{ab}$ ,  $C^{a0}$ ,  $C^{00}$

Operator	Definition
$C^{AB}$	$g^{\mu\nu} \partial_\mu \phi^A \partial_\nu \phi^B$ , $A, B = 0, 1, 2, 3$
$B^{ab}$	$C^{ab} = g^{\mu\nu} \partial_\mu \phi^a \partial_\nu \phi^b$ , $a, b = 1, 2, 3$
$Z^{ab}$	$C^{a0} C^{b0}$
$X$	$C^{00}$
$W^{ab}$	$B^{ab} - Z^{ab}/X$

# Lagrangian of 4 scalar + shift + $SO(3)_\phi$ symm.

## $SO(3)_\phi$ Scalars

$$L.V.: \quad u^\mu = -\frac{\epsilon^{\mu\alpha\beta\gamma}}{6 b \sqrt{-g}} \epsilon_{abc} \partial_\alpha \Phi^a \partial_\beta \Phi^b \partial_\gamma \Phi^c, \quad \mathcal{V}^\mu = -\frac{\nabla^\mu \Phi^0}{(-X)^{1/2}}$$

L.S.:

Operator	Definition
$X$	$C^{00}$
$b$	$\sqrt{\det \mathbf{B}}$
$Y$	$u^\mu \partial_\mu \Phi^0$
$y_n$	$\text{Tr}(\mathbf{B}^n \cdot \mathbf{Z}), \quad n = 0, 1, 2, 3$
$\tau_n$	$\text{Tr}(\mathbf{B}^n), \quad n = 1, 2, 3$
$w_n$	$\text{Tr}(\mathbf{W}^n), \quad n = 0, 1, 2, 3$
$\mathcal{O}_{\alpha\beta n}$	$(X/Y^2)^\alpha (y_n/Y^2)^\beta$

$$U_{SO(3)_\phi \text{ invariant}} = U(\underbrace{X, Y, y_n, \tau_n}_{9 \text{ operators}})$$

9 operators

# Shift Sym.+ $SO(3)_\phi$ + Non Linear Sym.

Four-dimensional media		
Symmetries of the action	LO scalar operators	Type of medium
$\Phi^A \rightarrow \Phi^A + f^A, \quad \partial_\mu f^A = 0$	$X, Y, \tau_n, y_n$	supersolids
$\Phi^a \rightarrow \Phi^a + f^a(\Phi^0)$	$X, w_n$	
$\Phi^0 \rightarrow \Phi^0 + f(\Phi^a)$	$Y, \tau_n$	
$\Phi^0 \rightarrow \Phi^0 + f(\Phi^0)$	$\tau_n, w_n, \mathcal{O}_{\alpha\beta n}$	
$\Phi^a \rightarrow \Phi^a + f^a(\Phi^0) \ \& \ \Phi^0 \rightarrow \Phi^0 + f(\Phi^0)$	$w_n$	
$V_S \text{Diff: } \Phi^a \rightarrow \Psi^a(\Phi^b), \quad \det  \partial\Psi^a/\partial\Phi^b  = 1$	$b, Y, X$	superfluids
$\Phi^0 \rightarrow \Phi^0 + f(\Phi^0) \ \& \ V_S \text{Diff}$	$b, \mathcal{O}_\alpha$	
$\Phi^0 \rightarrow \Phi^0 + f(\Phi^a) \ \& \ V_S \text{Diff}$	$b, Y$	perfect fluid
$\Phi^A \rightarrow \Psi^A(\Phi^B), \quad \det  \partial\Psi^A/\partial\Phi^B  = 1$	$b Y$	Superperfect fluid ( $\rho + p = 0$ )

# Operators with only $\Phi^0$ or $\Phi^a$ scalars + shift + $SO(3)_\phi$ + Non Linear Sym.

## Media with reduced internal dimensionality

Symmetries of the action	LO scalar operators	Type of medium
$(\Phi^a) \quad \Phi^a \rightarrow \Phi^a + c^a, \quad \partial_\mu c^a = 0$	$\tau_n$	solid
$(\Phi^a) \quad V_s \text{Diff}$	$b = b(\tau_n)$	perfect fluid
$(\Phi^0) \quad \Phi^0 \rightarrow \Phi^0 + c^0, \quad \partial_\mu c^0 = 0$	$X$	

$$b = \frac{1}{6}(2\tau_3 - 3\tau_2\tau_1 + \tau_1^3)$$



# Lagrangian Materials $\subset$ Massive Gravity Theories

It is possible to choose the comoving coord in such a way that the fluid system is at rest

Unitary Gauge:

If  $\text{Det}|\partial_\alpha \Phi^A(x)| \neq 0$ :  $\Phi^A(x) \equiv x^\mu \delta_\mu^A$

$\rightarrow$  Spontaneous Breaking of spacetime symmetries ( $T_X : x^\mu \rightarrow x^\mu + c^\mu$ )  
but Internal symmetry can compensate ( $T_\Phi : \Phi^A \rightarrow \Phi^A - c^A$ )

Residual symmetries:  $T_x + T_\Phi$  (the same for Rotations)

$\Rightarrow$  homogeneity and isotropy of background configurations

$$h_{AB} \partial_\mu \Phi^A \partial_\nu \Phi^B \rightarrow h_{\mu\nu} \quad \text{frozen metric}$$

(ex:  $\eta_{\mu\nu}$  (Lorentz Inv.),  $\delta_\mu^i \delta_{ij} \delta_\nu^j$ ,  $\delta_\mu^0 \delta_{00} \delta_\nu^0$  (Lorentz Br.) )

<b>LO self-gravitating media</b> $\mathcal{L}(C^{AB}, g_{\mu\nu})$ $\mathcal{O}_{LO} : X, Y, \tau_n, y_n$	Map Unitary gauge $\rightarrow$ $\leftarrow$ Stuckelberg trick	<b>Massive gravity</b> $\mathcal{L}(h_{\mu\nu}, g^{\mu\nu})$ ADM: $N, N^i, \gamma_{ij}$
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## Postulated Fundamental Relation

$s = s(\rho, n)$  or in the energy representation  $\rho = \rho(s, n)$

## I Law of Thermodynamics

$$\boxed{d\rho = T ds + \mu dn} \rightarrow T \equiv \left. \frac{\partial \rho}{\partial s} \right|_n, \quad \mu \equiv \left. \frac{\partial \rho}{\partial n} \right|_s.$$

## Euler relation

$$\boxed{\rho + p = T s + \mu n},$$

## Gibbs-Duhem relation

$$\boxed{dp = s dT + n d\mu}$$

# Hydrodynamical Equation of Motion

Hydrodynamic equations express conservation laws

- Diff. invariance:

$$\nabla_{\nu} T^{\mu\nu} = 0$$

Space time translations. Rotations and boost are identically conserved.

- Global U(1) symmetry:

$$\nabla_{\mu} n^{\mu} = 0 ;$$

variables:  $u^\mu$  ( $u^2 = -1$ ),  $n$ ,  $\rho$ ,  $p$

$$T_{\mu\nu} = \underbrace{(\rho + p)}_{\text{Euler: } T s + \mu n} u_\mu u_\nu + p g_{\mu\nu} = (T s_\mu + \mu n_\mu) u_\nu + p g_{\mu\nu}$$

Euler:  $T s + \mu n$

$$n^\mu = n u^\mu, \quad s^\mu = s u^\mu$$

$$\text{EoM} : \dot{\rho} + \theta(\rho + p) = 0 \rightarrow \underbrace{T \nabla^\alpha s_\alpha + \mu \nabla^\alpha n_\alpha}_{0 \leftrightarrow 0} = 0,$$

$$(\rho + p)\dot{u}_\mu + (\delta_\mu^\nu + u_\mu u^\nu)\nabla_\nu p = 0$$

$$\nabla^\alpha n_\alpha = 0 \Rightarrow \frac{\dot{s}}{s} = \frac{\dot{n}}{n} = -\theta \Rightarrow u^\mu \nabla_\mu \left( \frac{s}{n} \right) \equiv \dot{\sigma} = 0;$$

( $\theta = \nabla_\mu u^\mu$ ,  $\dot{f} = u^\mu \nabla_\mu f$ ) entropy per particle  $\sigma = s/n$  is conserved along the flow lines (adiabatic fluid).

# Perfect Fluids Lagrangian

$V_S$ Diff:  $\Phi^a \rightarrow \Psi^a(\Phi^b)$ ,  $\det |\partial\Psi^a/\partial\Phi^b| = 1$

$$S_{Fluid} = \int d^4x \sqrt{-g} U(\underbrace{b}_{\Phi^0 \rightarrow \Phi^0 + f(\Phi^a)}, \underline{Y}; \underbrace{X}_{\Phi^a \rightarrow 0})$$

$$u^\mu = -\frac{\epsilon^{\mu\alpha\beta\gamma}}{6 b \sqrt{-g}} \epsilon_{abc} \partial_\alpha \Phi^a \partial_\beta \Phi^b \partial_\gamma \Phi^c, \quad \mathcal{V}^\mu = -\frac{\nabla^\mu \Phi^0}{(-X)^{1/2}}$$

$$\underline{b} = \det |\partial_\mu \Phi^a \partial^\mu \Phi^b|^{1/2}, \quad \underline{Y} = u^\mu \partial_\mu \Phi^0, \quad \underline{X} = \partial_\mu \Phi^0 \partial^\mu \Phi^0$$

$$U(\underline{b}, \underline{Y}) \rightarrow T_{\mu\nu} = (U - b U_b) g_{\mu\nu} + (Y U_Y - b U_b) u_\mu u_\nu$$

$$U(\underline{X}) \rightarrow T_{\mu\nu} = U g_{\mu\nu} + 2 X U_X \mathcal{V}_\mu \mathcal{V}_\nu$$

$$T_{\mu\nu}^{Perfect Fluid} = \rho g_{\mu\nu} + (\rho + p) v_\mu v_\nu$$

Density, pressure and conserved currents for Perfect fluid Lagrangians

Lagrangian	$\rho$	$p$	Cons. Currents
$U(b)$	$-U$	$U - b U_b$	$J^\mu = b u^\mu$
$U(Y)$	$-U + Y U_Y$	$U$	$J_{(1)}^\mu = U_Y u^\mu$
$U(X)$	$-U + 2X U_X$	$U$	$J_{(2)}^\mu = -2(-X)^{1/2} U_X v^\mu$
$U(b, Y)$	$-U + Y U_Y$	$U - b U_b$	$J^\mu = b u^\mu, J_{(1)}^\mu = U_Y u^\mu$

# Matching Thermodynamics - Perfect Fluid Lagrangian

$\rho(\mathcal{I}, \mathcal{D}), p(\mathcal{I}, \mathcal{D})$      $2 \text{ indep. } \mathcal{I} + 2 \text{ dep. } \mathcal{D}(\mathcal{I})$

$$\underbrace{(\underbrace{\rho, p}_{\text{Thermo variables}}, \underbrace{s, T, n, \mu}_{\text{Thermo variables}})}_{\text{Thermo variables}} \Leftrightarrow U(\underbrace{\mathcal{O}}_{b, Y, X, (b, Y)}), T_{\mu\nu} \rightarrow p(\mathcal{O}), \rho(\mathcal{O})$$

ex:  $\mathcal{I} = (n, s)$ ,  $\mathcal{D} = (T, \mu)$ ,  $\mathcal{O} = b$  then I promote  $\mathcal{O} \rightarrow \mathcal{O}(\mathcal{I})$

$$U(b) : \boxed{\rho(n, s) = p(b) = -U(b)}, \quad p = U - b U_b \text{ with } b = b(n, s):$$

$b_s \equiv \partial_s b|_n, \quad b_n = \partial_n b|_s$

$$\rho = \rho(n, s) : \quad T(n, s) \equiv \left. \frac{\partial \rho}{\partial s} \right|_n = -U_b b_s,$$

$$\mu(n, s) \equiv \left. \frac{\partial \rho}{\partial n} \right|_s = -U_b b_n$$

Euler :  $\rho + p = s T + n \mu$  (It has to be valid for any  $U$ )

$$\Rightarrow -b U_b = -s U_b b_s - n U_b b_n \rightarrow U_b (b - s b_s - n b_n) = 0$$

$$b = s f\left(\frac{n}{s}\right) \text{ and } T = -U_b \left(f - \frac{n}{s} f'\right), \quad \mu = -U_b f'$$

# Matching Thermodynamics - Perfect Fluid Lagrangian

$$\rho(\mathcal{I}, \mathcal{D}), p(\mathcal{I}, \mathcal{D}) \quad 2 \text{ indep. } \mathcal{I} + 2 \text{ dep. } \mathcal{D}(\mathcal{I})$$

$$\underbrace{(\overbrace{\rho, p}, \overbrace{s, T, n, \mu})}_{\text{Thermo variables}} \Leftrightarrow U(\underbrace{\mathcal{O}}), T_{\mu\nu} \rightarrow p(\mathcal{O}), \rho(\mathcal{O})$$

$$b, Y, X, (b, Y)$$

We postulate  $\mathcal{O} = \mathcal{O}(\mathcal{I})$  such that: It has to be valid for any  $U$

$$\rho(\mathcal{I}, \mathcal{D}(\mathcal{I})) = \rho(\mathcal{O}(\mathcal{I})), \quad p(\mathcal{I}, \mathcal{D}(\mathcal{I})) = p(\mathcal{O}(\mathcal{I}))$$

$$\underbrace{\text{Law of Th.}}_{d\rho = Tds + \mu dn}, \quad \underbrace{\text{Euler eq.}}_{\rho + p = Ts + \mu n}, \quad \underbrace{\text{Gibbs Duhem}}_{dp = sdT + nd\mu} \Leftrightarrow \text{Conservation Law : } \nabla T = 0$$

$$\text{es: } \mathcal{I} = (n, T) \rightarrow \mathcal{D}(\mathcal{I}) = (\mu(n, T), s(n, T)) \text{ and } \mathcal{O} = (b, Y), \quad p = U - b U_b, \quad \rho = -U + Y U_Y$$

$$d\rho(\mathcal{O}(\mathcal{I})) = Tds(\mathcal{I}) + \mu(\mathcal{I})dn$$

$$\rho(\mathcal{O}(\mathcal{I})) + p(\mathcal{O}(\mathcal{I})) = T s(\mathcal{O}(\mathcal{I})) + \mu n(\mathcal{O}(\mathcal{I}))$$

$$b = n, \quad Y = T, \quad s = U_Y, \quad \mu = -U_b$$



# Thermodynamical potentials and related thermodynamical variables

Thermo Pot.	$\mathcal{I}$	Leg. Transf.	$\mathcal{D}$
energy $\rho$	$s, n$	none	$T = \left. \frac{\partial \rho}{\partial s} \right _n, \mu = \left. \frac{\partial \rho}{\partial n} \right _s$
free energy $\mathcal{F}$	$T, n$	$\mathcal{F} = \rho - T s$	$s = - \left. \frac{\partial \mathcal{F}}{\partial T} \right _n, \mu = \left. \frac{\partial \mathcal{F}}{\partial n} \right _s$
grand potential $\omega$	$T, \mu$	$\omega = -\rho$	$s = - \left. \frac{\partial \omega}{\partial T} \right _\mu, n = - \left. \frac{\partial \omega}{\partial \mu} \right _T$
potential $\mathcal{I}$	$s, \mu$	$\mathcal{I} = \rho - \mu n$	$T = \left. \frac{\partial \mathcal{I}}{\partial s} \right _\mu, n = - \left. \frac{\partial \mathcal{I}}{\partial \mu} \right _s$

The Thermodynamical connection with Perfect Fluid Lagrangian it works when

$$U \equiv - \textit{Thermodynamical Potential}$$

# Dictionary: Thermo $\leftrightarrow$ Lagrangian

Potentials	$U(b)$	$U(Y)$	$U(X)$	$U(b, Y)$
Indep. Variables	$\rho = -U$ $p = U - b U_b$ $J_\mu = b u_\mu$	$\rho = -U + Y U_Y$ $p = U$ $J_\mu = U_Y u_\mu$	$\rho = -U + 2X U_X$ $p = U$ $J_\mu = -2(-X)^{1/2} U_X v_\mu$	$\rho = -U + Y U_Y$ $p = U - b U_b$ $J_\mu^{(1)} = b u_\mu$ $J_\mu^{(2)} = U_Y u_\mu$
$(\mu, s)$ $\mathcal{I} = -U$	$b = s$ $n = 0$ $T = -U_b$	$Y = \mu$ $n = U_Y$ $T = 0$	$X = -\mu^2$ $n = -2 U_X \sqrt{-X}$ $T = 0$	$b = s, Y = \mu$ $n = U_Y$ $T = -U_b$
$(n, T)$ $\mathcal{F} = -U$	$b = n$ $s = 0$ $\mu = -U_b$	$Y = T$ $s = U_Y$ $\mu = 0$	$X = -T^2$ $s = -2 U_X \sqrt{-X}$ $\mu = 0$	$b = n, Y = T$ $s = U_Y$ $\mu = -U_b$
$(\mu, T)$ $\omega = -U$		$Y = T f\left(\frac{\mu}{T}\right)$ $n = U_Y f'$ $s = U_Y\left(f - \frac{\mu}{T} f'\right)$	$X = -\mu^2 f\left(\frac{T}{\mu}\right)$ $n = U_X(T f' - 2\mu f)$ $s = -U_X \mu f'$	
$(n, s)$ $\rho = -U$	$b = s f\left(\frac{n}{s}\right)$ $\mu = -U_b f'$ $T = -U_b\left(f - \frac{n}{s} f'\right)$			

$$b \rightarrow n, s, \quad Y \rightarrow \mu, T, \quad X \rightarrow \mu, T$$

# Barotropic Fluids $p = p(\rho)$

The perfect fluids  $U(b)$ ,  $U(Y)$ ,  $U(X)$  are all barotropic.

$p = U - b U_b$ ,  $\rho = -U$ ;  $p = f[\rho] \rightarrow$  Differential eq for  $\rho(b) : -\rho + b \rho_b = f[\rho]$

$p = w \rho$  with constant  $w$

$$U(b) = b^{1+w}, \quad U(Y) = Y^{(1+w)/w}, \quad U(X) = X^{(1+w)/2w},$$

**Cold Dark Matter** ( $w = 0$ ):  $\rho = m n \rightarrow U = -m b$  with  $b = n$ ,  $s = 0$ ,  $\mu = m$

**Radiation** ( $w = \frac{1}{3}$ ):  $\rho = 3 p = T^4$ ,  $U = Y^4/3$  with  $Y = T$ ,  $s = Y^3$ ,  $\mu = 0$

**Chaplygin gas**:  $p = -A/\rho$ , with constant  $A$

$$U(b) = \sqrt{A + \lambda b^2}, \quad U(Y) = \sqrt{A + \lambda Y^2}, \quad U(X) = \sqrt{A + \lambda X}$$

# Bose-Einstein and Fermi-Dirac distributions

Relativistic limits:  $T \gg m$ ,  $\alpha = e^{\frac{\mu}{T}}$

$$\underbrace{p = \frac{\rho}{3}}_{\text{barotropic eq of state}} = -\epsilon \frac{g}{\pi^2} T^4 \text{Li}_4(-\epsilon \alpha),$$

*barotropic eq of state*

$$n = -\epsilon \frac{g}{\pi^2} T^3 \text{Li}_3(-\epsilon \alpha), \quad s = \frac{4}{3} \frac{\rho}{T} - n \log(\alpha);$$

$$\boxed{U(Y) = \frac{g}{3} Y^4}, \quad \boxed{Y = T f\left(\frac{\mu}{T}\right)} \text{ where } f\left(\frac{\mu}{T}\right) = \left[-\frac{1}{\pi^2} \epsilon \text{Li}_4\left(-\epsilon e^{\frac{\mu}{T}}\right)\right]^{1/4}$$

For  $\frac{\mu}{T} \ll 1$  we have  $\underline{Y \rightarrow T}$ ; for  $\frac{\mu}{T} \gg 1$  we have  $\underline{Y \rightarrow \mu}$

# Bose-Einstein and Fermi-Dirac distributions

Non relativistic limit:  $m \gg T$

$$n = g \left( \frac{m T}{2\pi} \right)^{3/2} e^{\frac{(\mu-m)}{T}}, \quad \rho = n \left( m + \frac{3T}{2} \right), \quad p = n T \ll \rho.$$

$$b=n, Y=T$$

$$\underline{U(b, Y)} = b Y \left\{ 1 + \text{Log} \left[ \frac{g}{b} \left( \frac{m Y}{2\pi} \right)^{3/2} \right] \right\} - b m.$$

$$(p + \kappa n^2)(1 - \gamma n) = T n,$$

$$b=n, Y=T$$

$$U(b, Y) = b \left\{ b\kappa + Y - m + Y \log \left[ \frac{g(1 - \gamma b)}{b} \left( \frac{m Y}{2\pi} \right)^{3/2} \right] \right\}.$$

$$\rho = b \left( m + \frac{3}{2} T - n\kappa \right);$$

$$s = \frac{5b}{2} + b \log \left[ \left( \frac{m Y}{2\pi} \right)^{3/2} \frac{g(1 - b\gamma)}{b} \right];$$

$$\mu = m - 2b\kappa + \frac{bY\gamma}{1 - b\gamma} + Y \log \left[ \frac{b}{g(1 - b\gamma)} \left( \frac{m Y}{2\pi} \right)^{-3/2} \right].$$

# Thermodynamical Stability

The entropy density  $s = s(\rho, n)$  is a convex function, thus the Hessian matrix  $|s_{ij}| \equiv |\partial_{ij}^2 s|$ ,  $i, j = \rho, n$ , is negative definite

$$\underline{s_{\rho\rho} + s_{nn} \leq 0}, \quad \underline{s_{\rho\rho} s_{nn} - s_{\rho n}^2 \geq 0}$$

$$\omega_{\mu\mu} + \omega_{TT} \leq 0, \quad \omega_{\mu\mu} \omega_{TT} - \omega_{\mu T}^2 \geq 0 \quad \underline{\omega = -U(X), U(Y)}$$

$$U_X + 2X U_{XX} \leq 0, \quad (f'^2 - 2f f'') U_X \geq 0, \quad X = -T^2 f[z], \quad z = \frac{\mu}{T},$$

$$U_{YY} \geq 0, \quad f'' U_Y \geq 0, \quad Y = T f[z], \quad z = \frac{\mu}{T},$$

$$\rho_{ss} + \rho_{nn} \geq 0, \quad \rho_{ss} \rho_{nn} - \rho_{sn}^2 \geq 0 \quad \underline{\rho = -U(b)}$$

$$U_{bb} \leq 0, \quad f'' U_b \leq 0, \quad b = s f[z], \quad z = \frac{n}{s},$$

$$\mathcal{I}_{\mu\mu} \leq 0, \quad \mathcal{I}_{ss} \geq 0 \quad \underline{\mathcal{I} = -U(b, Y)}$$

$$U_{bb} \leq 0, \quad U_{YY} \geq 0, \quad b = n, s, \quad Y = T, \mu,$$

for  $f = f_{B/FD}$  we have  $f'' \geq 0$  and  $(f'^2 - 2f f'') \leq 0$  so that

$$U_{YY}, U_Y \geq 0; \quad U_{bb}, U_b \leq 0; \quad U_X + 2X U_{XX}, U_X \leq 0$$

# Goldstone Bosons fluctuations

$$\Phi^i(t, x^j) = x^i + \pi^i(t, x^j), \quad \Phi^0 = t + \pi^0$$

**Thermo stability** +  $(\rho + p) \geq 0$   $\Rightarrow$  **Dynamic stability** but Th.s.  $\Leftrightarrow (\rho + p) \geq 0 + D.s.$

$$S^{(2)}[b] = \frac{1}{2}(\rho + p) \int d^4x \left( \dot{\pi}^i \dot{\pi}^i - c_s^2 (\partial_i \pi^i)^2 \right),$$

$$U_{bb} \leq 0, f'' U_b \leq 0 \quad \leftrightarrow \quad (\rho + p) = -b U_b \geq 0, c_s^2 = \frac{b U_{bb}}{U_b} \geq 0$$

$$S^{(2)}[\gamma] = \frac{1}{2}(\rho + p) \int d^4x \left( \dot{\pi}^i \dot{\pi}^i - c_s^2 (\partial_i \pi^i)^2 \right)$$

$$U_{\gamma\gamma} \geq 0, f'' U_\gamma \geq 0 \quad \leftrightarrow \quad (\rho + p) = \gamma U_\gamma \geq 0, c_s^2 = \frac{U_\gamma}{\gamma U_{\gamma\gamma}} \geq 0$$

$$S^{(2)}[X] = \frac{1}{2}(\rho + p) \int d^4x \left( c_s^{-2} (\dot{\pi}^0)^2 - (\partial_i \pi^0)^2 \right)$$

$$2XU_{XX} + U_X \leq 0, ((f')^2 - 2ff'') U_X \geq 0 \quad \leftrightarrow \quad (\rho + p) = 2XU_X \geq 0, c_s^2 = \frac{U_X}{2XU_{XX} + U_X} \geq 0$$

$$S^{(2)}[b, \gamma] = \frac{1}{2} \int d^4x \left[ (\rho + p) \dot{\pi}^i \dot{\pi}^i + \rho_\gamma (\dot{\pi}^0)^2 - p_b (\partial_i \pi^i)^2 - 2c_b^2 \rho_\gamma \dot{\pi}^0 \partial_i \pi^i \right]$$

$$U_{bb} \leq 0, U_{\gamma\gamma} \geq 0, \quad (\rho + p) = \gamma U_\gamma - b U_b \geq 0, c_s^2 = \frac{(b U_{b\gamma} + U_\gamma)^2 - b^2 U_{bb} U_{\gamma\gamma}}{U_\gamma^2 (\rho + p)} \geq 0$$



$$c_s^2 \equiv \frac{\dot{p}}{\dot{\rho}}$$

For barotropic perfect fluids  $U(b)$ ,  $U(Y)$ ,  $U(X)$

$$\delta p = c_s^2 \delta \rho \rightarrow c_s^2 = \frac{dp}{d\rho} : \frac{b U_{bb}}{U_b}, \frac{U_Y}{Y U_{YY}}, \frac{U_X}{U_X + 2XU_{XX}}$$

$U(b, Y)$  is not barotropic

$$\delta p = \underbrace{c_s^2 \delta \rho}_{\text{adiabatic modes}} + \underbrace{b Y (c_b^2 - c_s^2) \delta \sigma}_{\text{entropic modes}}, \quad c_b^2 = \left. \frac{\partial p}{\partial \rho} \right|_b, \quad \sigma = \frac{s}{n} \text{ with } \dot{\sigma} = 0$$

- Media classification and internal scalar symmetries
- Media Lagrangians  $\subset$  Massive Gravity
- Perfect Fluid Lagrangians
- Thermodynamics of Perfect Fluid Lagrangians
- Thermodynamical and Dynamical stability

# Non perturbative # DoF and structure of the Potentials

ADM variables:  $g^{\mu\nu} = \begin{pmatrix} -N^{-2} & \\ & \gamma^{ij} - N^{-2} N^i N^j \end{pmatrix}$ , ( $N^A \equiv (N, N^i)$ ,  $\gamma^{ij}$ )

- **6 DoF**:  $U = U[N^A, \gamma]$ : Supersolid
- **5 DoF**:  $U = (U[\mathcal{K}] + (\mathcal{E}[\xi, \gamma] + \mathcal{U}_{\xi^i} Q^i)/N)$

where  $N^i = N \xi^i + Q^i$ ,  $\mathcal{K}^{ij} = \gamma^{ij} - \xi^i \xi^j$  and  $Q^i = (\partial_{\xi^i}^2 U)^{-1} \partial_{\xi^j} \mathcal{E}$

Solids  $U(\tau_n)$  and Perfect fluids  $U(b = b(\tau_n))$ ,  $\mathcal{E} = 0$

- **3 DoF**:  $U = U[N, \gamma]$  Perfect fluid  $U(X)$
- **3 DoF**:  $U = (\mathbf{U}[\gamma] + \mathbf{E}[\gamma]/N)$
- **2 DoF**:  $U = (\Lambda + \mathbf{E}[\gamma]/N)$

$$N \rightarrow \boxed{X = -\frac{1}{N^2}}, \quad \gamma \rightarrow \boxed{w_n = \text{Tr}[\gamma^n]}, \quad \mathcal{K} \rightarrow \boxed{\tau_n = \text{Tr}[(\gamma - \xi \otimes \xi)^n]},$$

$$Y = \frac{1}{N \sqrt{1 - \xi^2}}, \quad b = \sqrt{(1 - \xi^2) \gamma}, \quad y_n = \frac{\xi \cdot (\gamma - \xi \otimes \xi) \cdot \xi}{N^2}, \quad O_{\alpha\beta n} = (1 - \xi^2)^{\alpha+\beta} \xi \cdot (\gamma - \xi \otimes \xi) \cdot \xi$$

$$S_{Solids} = \int d^4x \sqrt{-g} U(\tau_{n=1,2,3} = \text{Tr}[(B \equiv \partial\Phi^a \partial\Phi^b)^n])$$

Hook Potential

$$U(\tau_i) = \sqrt{\det|B|} (\lambda \text{Tr}[I - B]^2 + 2 \mu \text{Tr}[(I - B)^2]) =$$

$$b (6 \mu + 9 \lambda - (6 \lambda + 4 \mu) \tau_1 + \lambda \tau_1^2 + 2 \mu \tau_2)$$

$\lambda \mu$ : Lamé' Constants,  $B$  strain tensor

$$T_{\mu\nu} = \underbrace{U}_{\rho} u_\mu u_\nu + \underbrace{2 b \partial_\mu \Phi^a ((-(6\lambda + 4\mu) + 2\lambda\tau_1)\delta_{ab} + 4\mu B_{ab}) \partial_\nu \Phi^b}_{\rho g_{\mu\nu} + \Pi_{\mu\nu}}$$